The Two Micron All Sky Survey

Roc Cutri and the 2MASS Team - November 2005

http://ipac.caltech.edu/2mass/ Skrutskie et al. 2006, AJ, 131, in press







2MASS Objectives

Technical Objective: To carry out a highly uniform, digital imaging survey of the entire sky in three near infrared bands

Science Goal: Produce an Image Atlas and Catalogs that will enable astronomers to:

 Trace luminous matter over the entire length of the Milky Way, probing largest structures of the Galaxy

– Construct the first all-sky photometric census of galaxies in the local universe to establish the foundations for measuring 3-D distribution of matter and mapping the Hubble flow out to $z\sim0.1$

 Have the statistical basis to search for rare but astrophysically important objects such as low mass stars/brown in the Solar neighborhood, globular clusters in the Galactic Plane and dust-obscured active galactic nuclei



Implementation

- Joint project of the Univ. of Massachusetts and IPAC/Caltech, funded primarily by NASA and NSF
- Identical, dedicated 1.3 m telescopes at Mt. Hopkins and CTIO
- Cameras containing three NICMOS 3 arrays for simultaneous imaging in:
 - J (1.24μm), H (1.66μm) and K_s (2.16μm)
- 2"/pixel 8.5'x8.5' FOV
- Highly automated operation using prescripted nightly plan





Survey Strategy

- Sky divided into 59,650 Tiles (6°x8.5')
- High efficiency "freeze-frame" scanning
 - Telescope scans in declination (1°/min)
 - Secondary freezes sky on focal planes for 1.5s
 - Two readouts at each position (0.05s & 1.3s)
 - Advance 1/6 frame (net 7.8s on each sky point)
 - Arrays rotated for sub-pixel sampling
- One scan yields 273 frames x 2 reads x 3 bands
- 65 Tiles (~60 deg²) scanned each average clear night



Sequence of 273 2MASS J-band frames from **Scan** of Abell 3627 region



Survey Operations

- Mt. Hopkins
 - June 1997 December 2000
 - 42% photometric
- CTIO
 - March 1998 February 2001
 - 64% photometric
- 24.5 TB raw data collected
- Photometric data covering
 99.998% of the sky

- 0.7 deg² in 21 small gaps



Survey Sky Coverage



- ~30% of the sky observed more than once
- Catalogs use only one epoch for uniformity
- Multi-epoch data to be released with *Extended Mission* products



Photometric Calibration

11111n AP ZERO POINTS J-CHANNEL: PHOTOMETRIC [CAL STRAT=2]

- Hourly observations of one of 35 (8.5'x 1°) Calibration Fields
 - Scanned 6 times each visit
 - 5-8 different Tiles/night
 - 5-50 standard stars per field
- Nightly photometric transformations
 - Zero point fit to nightly observations
 Atmospheric extinction derived from long-term statistics
- Photometric uniformity

 North/south observations tied via observations of the same equatorial fields

Internal consistency of standard star network (Nikolaev et al. 2000, AJ, 120, 3340)

 <1-2% global zero point variation, no systemmatic color drifts > 1%





Value of Touchstone Fields

- Cal fields observed 600-3700 times during survey sampling all possible observing conditions
- Instantaneous performance monitoring for quality assurance
- Empirical relationship between sensitivity and seeing, background, transparency, etc. applied to general survey measurements





Value of Touchstone Fields



Cartesian Sky Map of 2MASS Ks-band SNR=10 level (14.3 - 15.5 mag)



All-Sky Image Atlas

- 4,121,439 calibrated J, H, K_s images
- 512x1024 1"/pix FITS format
- 3-bands registered
- *Full-fidelity* Atlas Images available on-line via IRSA/2MASS Image Services

(http://irsa.ipac.caltech.edu/applications/2MASS/IM)



All-Sky Point Source Catalog (PSC)

 Photometry & astrometry for 471 million srcs (>95% MW stars)

• ~20 mags dynamic range (-4.4< K_s<15.5)

- Complete to **SNR>10** for J<15.8, H<15.1, K_s<14.3 mag
- Astrometric accuracy: ~80 mas wrt Tycho 2

All-Sky Extended Source Catalog (XSC)

2MASS Extended Sources Integrated Flux

Ks: 8.0–14.0 mag 18'/pixel

- Photometry, positions and shape info for **1.6 million resolved srcs** (~97% are z<0.1 galaxies)
- ~13.5 mags dynamic range (1.0<K_s<14.5)
- 92% Complete to K_s<13.5
- Positional accuracy ~0.8" wrt FIRST



Extinction Mapping Using Bulk Colors

IRAS 100µm





Tracing the Structure of the Milky Way



Skrutskie et al. 2001, BAAS, 33, 1437; Cole & Weinberg 2002, ApJ, 547, L43



Sagittarius Dwarf Tidal Stream



Majewski et al. 2003, ApJ, 599, 1082



Milky Way/Sagittarius Dwarf "fly-around" Animation by David Law/Univ. of Virginia



Foundation for L and T Dwarfs



- Prior to 1994, latest main sequence star known was M9.5
- DENIS, SDSS and 2MASS provided long wavelength sensitivity, depth and sky coverage to find rare, cooler objects
- New spectral types L & T
- Only problem is that they aren't rare!
- >400 L dwarfs and >60 T dwarfs now known

See new review by Kirkpatrick (2005 Ann.Rev.As.Ap.)

Structure of the Local Universe

Structure of the Local Universe

Large Scale Structure in the Local Universe





Galaxy Luminosity Function

- XSC + redshift surveys – ZCAT (Kochanek et al. 2001, ApJ, 560, 566) – 2dFGRS (Cole et al. 2001, MNRAS, 326, 255)
- Schechter function
 - $-M^{*}(K) = -23.4$
 - Late type more numerous
- M/L => mass function
 - M_{stars}^* = 3-7 x 10¹⁰ h⁻² M_o - Ω_{stars} = 1.6-2.9 x 10⁻³





Defining Characteristics

• Well-defined, concise set of performance requirements from beginning

 Sky coverage, photometric and astrometric accuracy, completeness, reliability, uniformity

- Prototyping
- Touchstone calibration fields to impose uniformity, monitor performance
- High throughput data processing system with extensive validation
- Staged data releases, with planned reprocessing ("You can't get it right the first time")
 - "Early" data releases take on a life of their own, though
- Well-characterized, understandable, documented data products
- Good access tools (bulk distribution, on-line via IRSA, CDS, etc)
- Legacy 2MASS data used in >1200 refereed publications