Cosmic Magnification with MSSS

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 Shear with Patel, Beswick & Muxlow (Merlin+VLA HDFN)

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This can eventually be attempted with E-LOFAR with long baselines - great way of finding out about the dark

Cosmic Magnification: Two competing effects











Selection being discussed; for simplicity here z>0.3

Expected N(>S)





 Correlate foreground (eg SDSS) and background (MSSS) object densities



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- Correlate foreground (eg SDSS) and background (MSSS) object densities
- The foreground objects are acting as lenses

• The anti-correlation is due to magnification; closely related to matter power spectrum



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Cosmology,
geometry
$$\xi_{\mu\delta}(\phi) = \frac{3H_0^2\Omega_0}{c^2} \int_0^{w_{\rm H}} dw' f_K(w') \bar{W}_Q(w') G_G(w') a^{-1}(w')$$
$$\times \int_0^{\infty} \frac{k \, dk}{2\pi} P_{\delta}(k, w') J_0[f_K(w')k\phi] .$$



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Cosmology, Background geometry sources $\xi_{\mu\delta}(\phi) = \frac{3H_0^2\Omega_0}{c^2} \int_0^{w_{\rm H}} dw' f_K(w') \bar{W}_Q(w') G_G(w') a^{-1}(w')$ $\times \int_0^\infty \frac{k \, dk}{2\pi} P_{\delta}(k, w') J_0[f_K(w')k\phi] .$



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Cosmology, Background
geometry sources Lenses

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Matter power spectrum

with Hoyle

Cosmic magnification predictions for MSSS



Cosmic magnification predictions for MSSS



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Issues

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- Need redshift distribution

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- One small step for radio weak lensing...