



A Method For Producing Matched Aperture Photometry Using Radial Profiles

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Surveys I

 As of DR8, 14,555 sq.
 Deg of u,g,r,i,z photometry

Depth of r_m~23.3

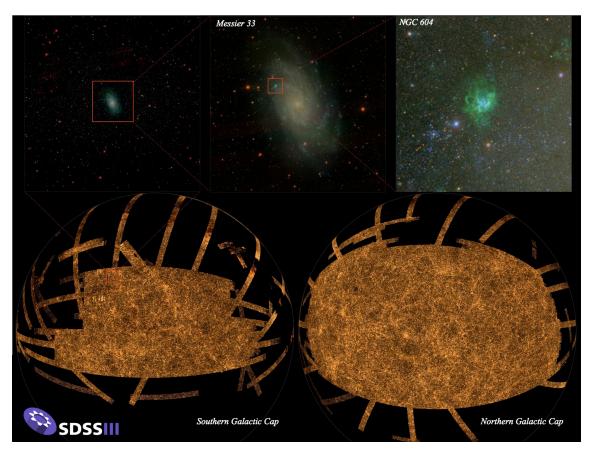


Image credit: SDSS3 Press release

SDSSIII

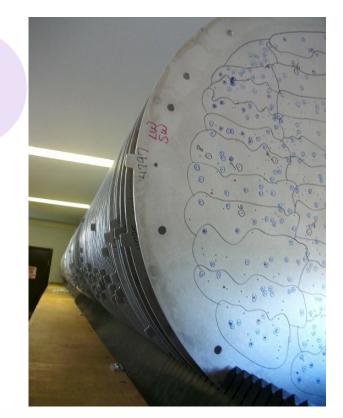


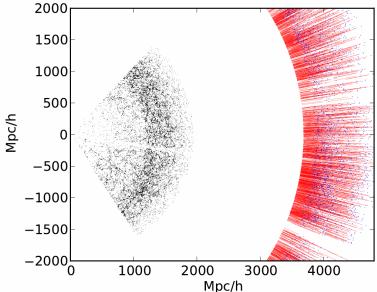
Surveys II Solution

BOSS

- ○1.5 million MassiveGalaxies at redshift 0.7
 - Also doing Lya Forest QSO measurements
- ○10,000 deg² of sky
- Provide % level
 cosmological constraints
- OHuge galaxy sample
 - Massive impact on galaxy evolution

Image credits: me (top), slozar et al 2011 (bottom)





Surveys III

AUS survey

- Redshift survey using AAOMega on AAO
- OThousands of LRGs at redshifts up to 1.2
- ~150 square degrees of equatorial fields

OUse of

AUS+BOSS/2SLAQ+SDSSII gives redshifts from 0.2—1.2

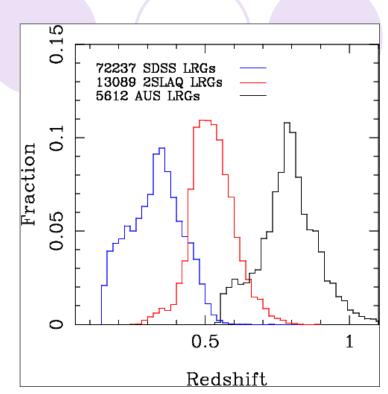


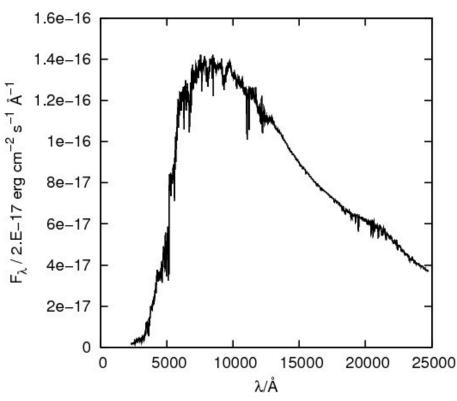


Image Credits: David Wake (top), David Malin (bottom)

Problems



- Spectroscopic surveys pushing to higher redshifts
 - OBy redshift 0.6, 4000Å break is in r band
 - OBy redshift 1, it is in z
 - Higher redshift galaxies don't have much observer's frame optical flux
- Degeneracies:
 - Template fitting has many parameters
 - OAge
 - ODust Reddening
 - OMetallicity
 - OSFR
- Get Degenerate Solutions
 NIR photometry helps

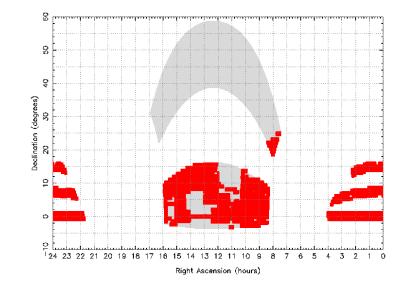


Maraston et al. 2009 LRG template



Surveys IV UKIDSS

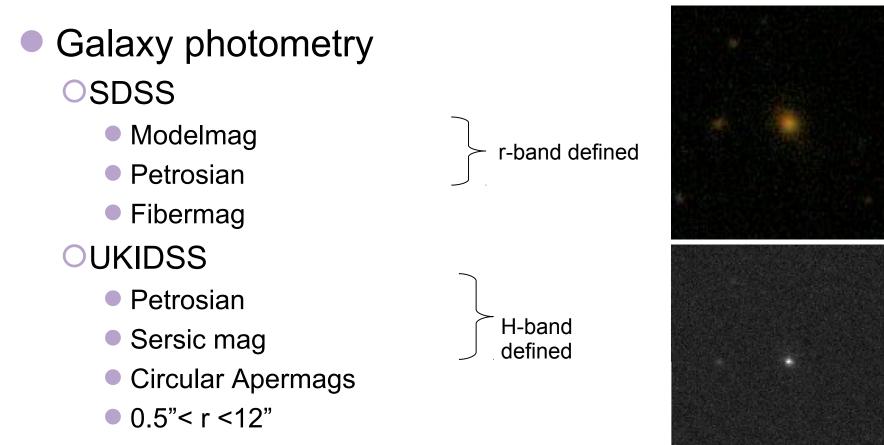
- UKIDSS Large Area Survey
- Y, J, H and K band photometry to depth of K=20.3
- Large overlap with BOSS, Aus
 - Get much better
 determination of galaxy
 SED







Photometric Catalogues



OMatched photometry hard from these!



Photometric Catalogues

Ideal solution is full re-extraction from images

- Oownload all imaging data
- ODegrade to standardised PSF

Integrate within same aperture across all bands

Hill et al 2010 did this for GAMA fields

○R-band defined Kron aperture photometry and 2" PSF

"Correct" method, expensive though

OTime

Computationally

- 4 hours to create 20GB mosaic
- "a few days" to create catalogue for each mosaic in dual image mode

Hill et al MNRAS 2010



Alternative Approach

Short cut for LRGs

- Small observed sizes
- Relatively simple properties
 - Radial profile
 - Stellar populations etc
- Can use alternative catalogue data to pre-derived magnitudes

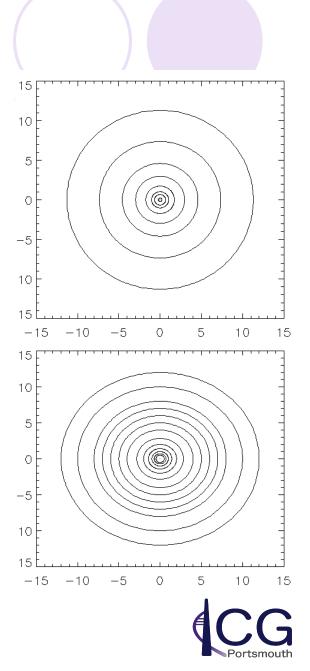
Use GAMA fields and compare to their matched photometry
 1300 SDSS+GAMA galaxies
 2700 BOSS+GAMA galaxies



Our Method

Obtain galaxy radial profile

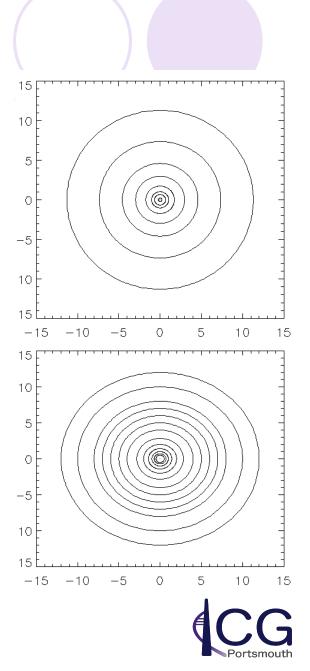
- SDSS photoprofile
 OFlux densities within annular bins
- UKIDSS Apermag photometry
 - Subtract off interior flux within each apermag
 - OGet flux densities within annular bins too



Our Method

Obtain galaxy radial profile

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Our Method II

- Take observed radial profile for each band
 Interpolate with cubic spline
 Project in 2D at a finer pixel scale
 Convolve mock image with gaussian to give
 - standardised seeing
 - OIntegrate within same aperture on all bands
 - Matched aperture photometry!

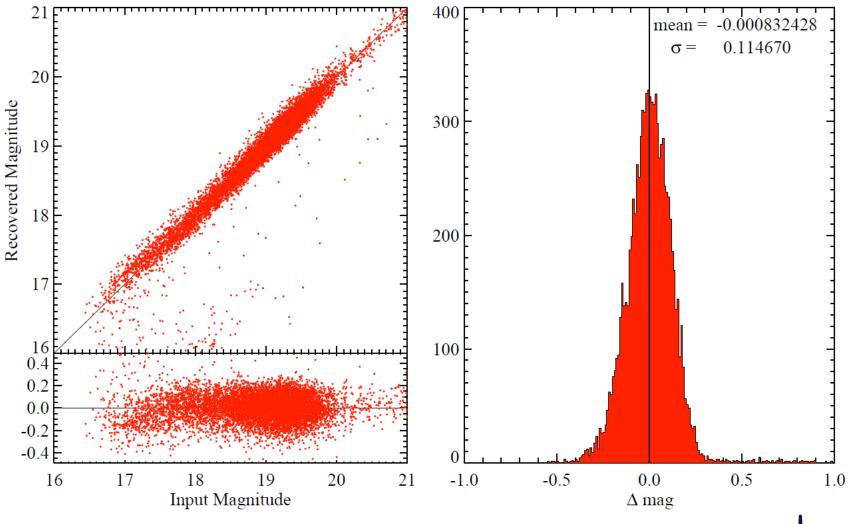


Tests

Simulated data created • Take pure De Vaucouleurs profile Convolve to SDSS-like seeing Take aperture photometry OMeasure photoprofile OAdd noise Feed into method Take aperture photometry

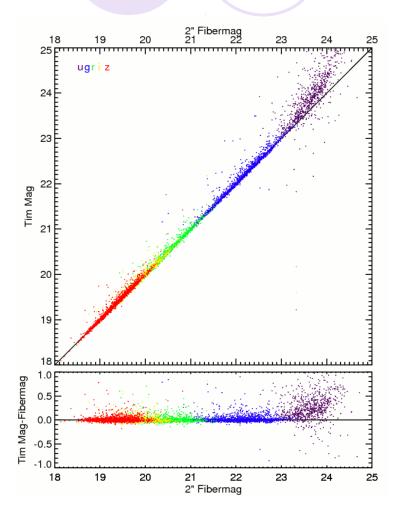


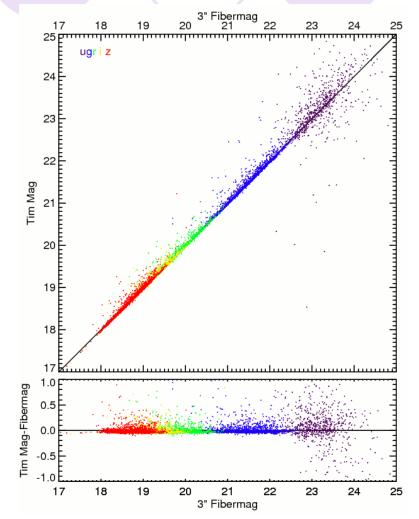
Results



CG

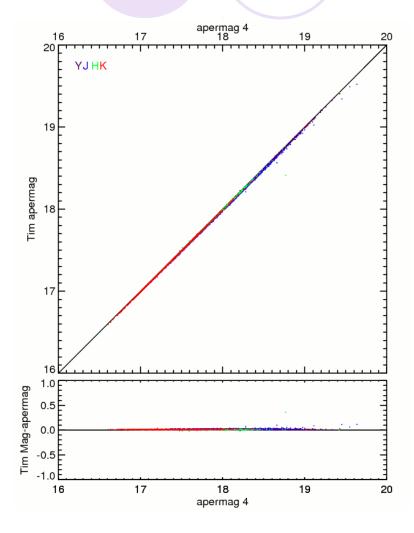
Results II – Reproducing Fibermags

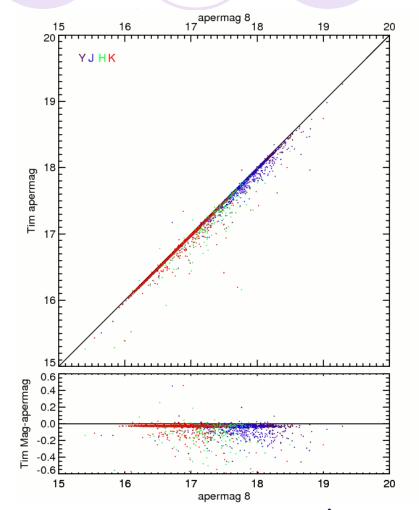




CG

Results III – Reproducing Apermags





Portsmouth

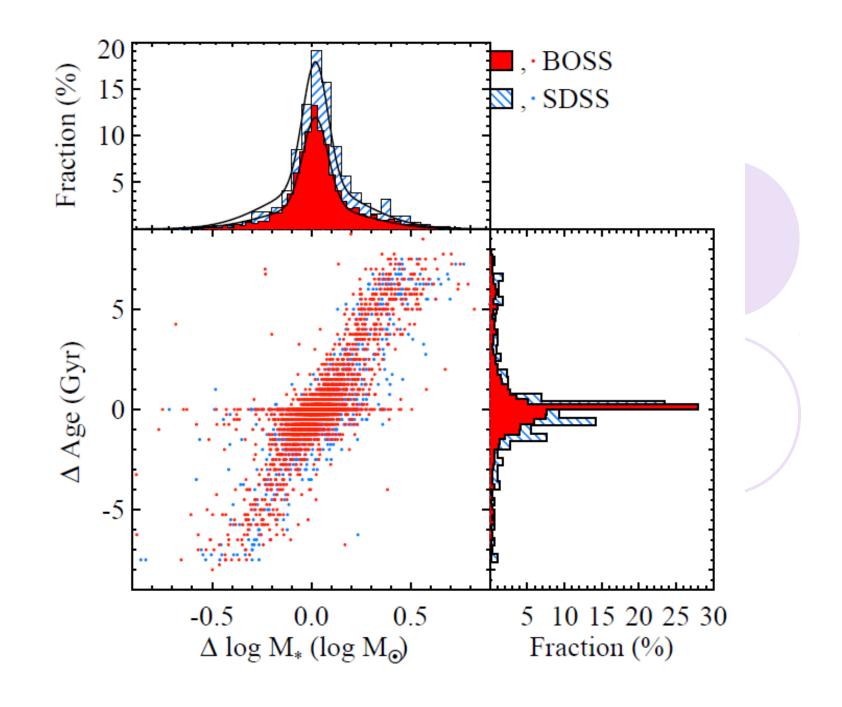
Tests – Template Fits

To test against GAMA, magnitudes dependent upon the aperture

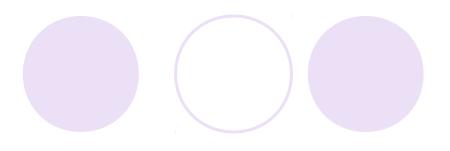
OInstead, use template fits

- Sensitive to colours and total flux
- Good test of matching
- Compare best fit template fits for GAMA and our photometry





Performance



Generally performs well

- OStellar masses and ages generally consistent
- Some age errors which then translate into stellar mass errors
 - Interpolated profile
 - Goes wrong sometimes
 - Colour error from upward bump in interpolation
 - Can remove these by comparison with catalog magnitudes' colours

 Takes 15 mins to do 9 bands of photometry for 2700 galaxies



Conclusions

- Can create a simple matched aperture SDSS-UKIDSS catalogue
- results looking good
 - OCan recreate simple apertures with good agreement
 - On average, a 20kpc radius aperture is consistent with elliptical Kron apertures
 - OStellar masses are close
 - Scattered around a zero offset
 - Some age errors causing outliers

