



THE SECOND XCS DATA RELEASE AND THE CLUSTERS IN THE SDSS AREA

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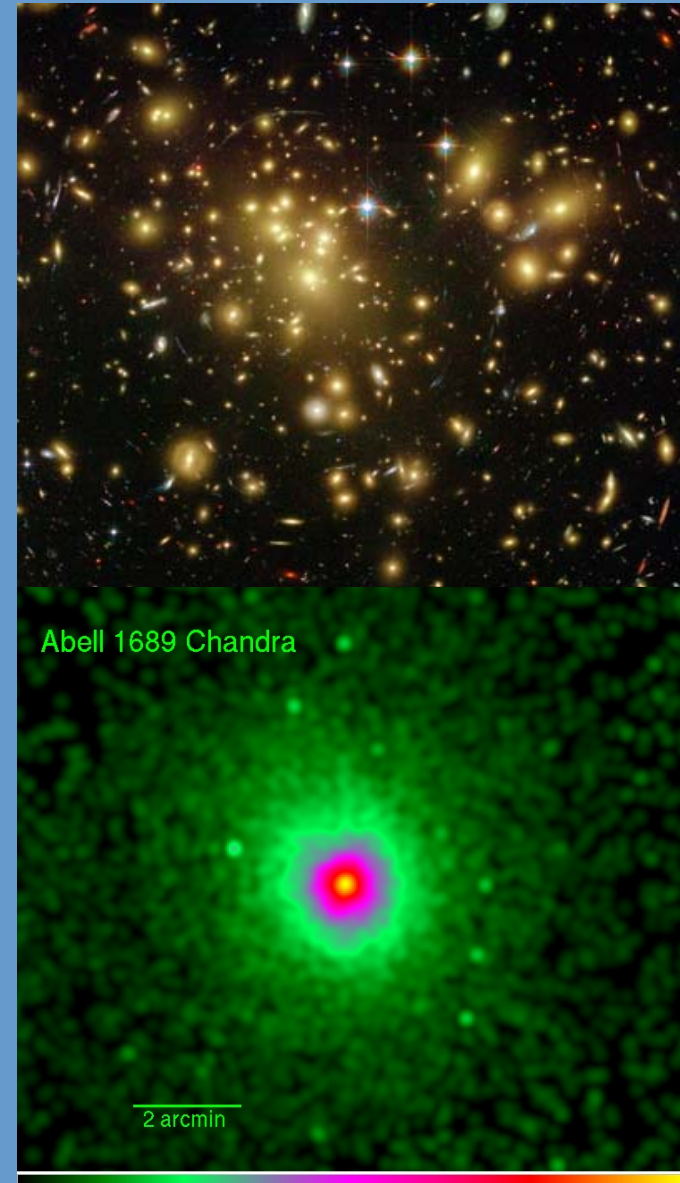
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Outline

- Galaxy clusters
- The XMM Cluster Survey (XCS)
- Optical confirmation
- Cluster redshifts
- Clusters in voids
- Future work

Galaxy clusters

- Largest gravitationally bound objects
- Dark matter – 90%
Hot gas – 9%
Galaxies – 1%
- Optical: red sequence, brightest cluster galaxy (BCG)
X-rays: bright, extended sources, intracluster gas emission
- - Cluster scaling relations (L_x , T_x , λ)
→ calculate cluster masses →
constrain cosmological parameters
(e.g. Allen, Evrard & Mantz 2011)
- Study cluster galaxy (e.g. BCG)
properties & evolution
- Need galaxy cluster catalogues!



The XMM Cluster Survey (XCS)

Serendipitous search for galaxy clusters in the public data of the XMM-Newton Science Archive across the whole sky

DR1: 503 optically confirmed X-ray clusters (Mehrtens et al 2012)

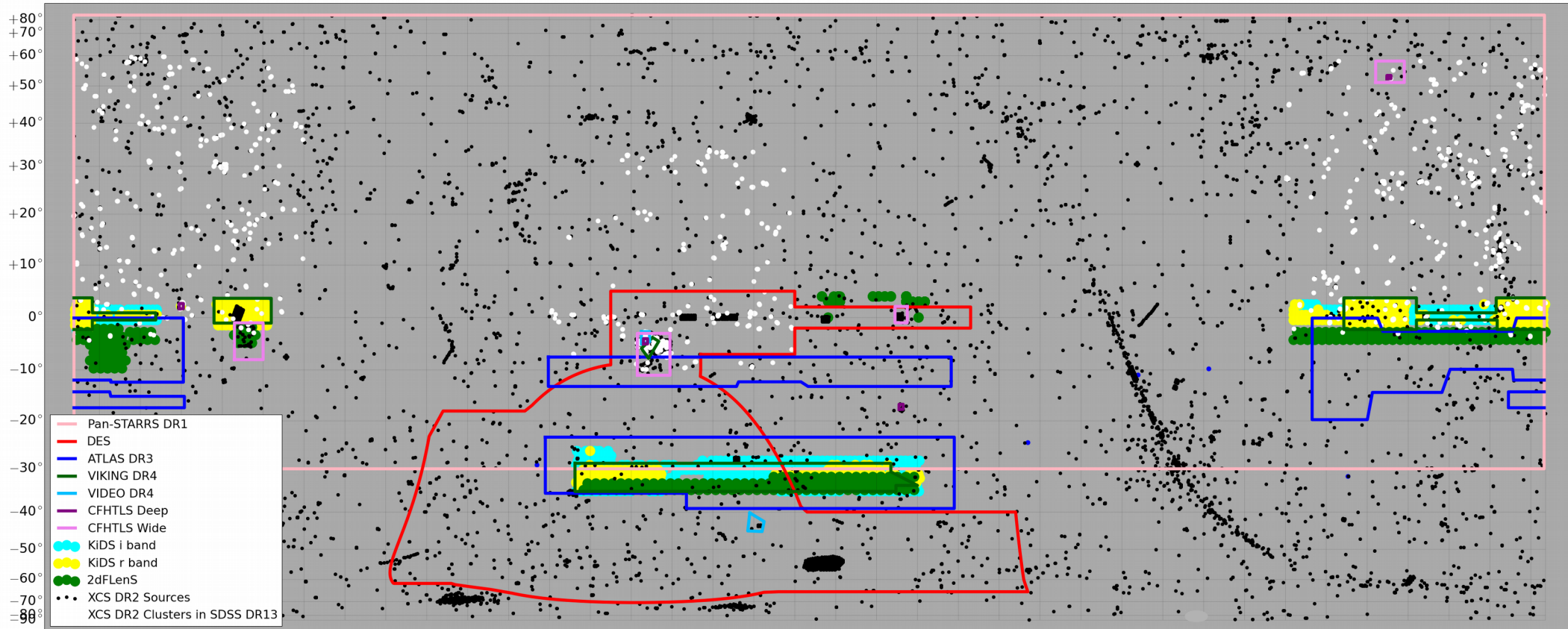
DR2: ~850 clusters just in SDSS DR13 area & improved pipelines for luminosities and temperatures

The largest galaxy cluster catalogue so far!!

	XXL	XCLASS	XMM-LSS	REFLEX
Area (deg ²)	2x25	All-sky	11	~14000
No. of clusters	450	850	52	452



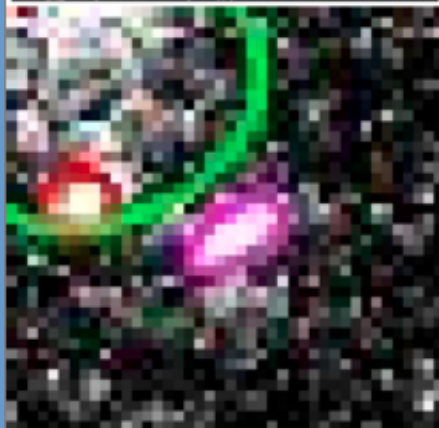
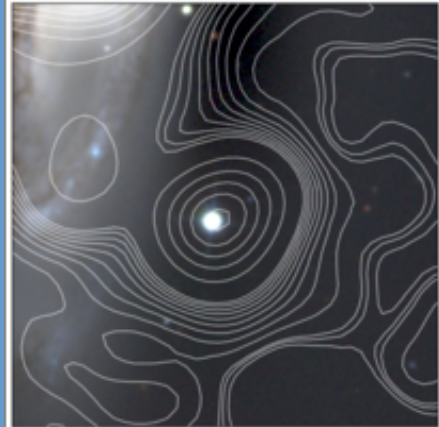
The XMM Cluster Survey (XCS)



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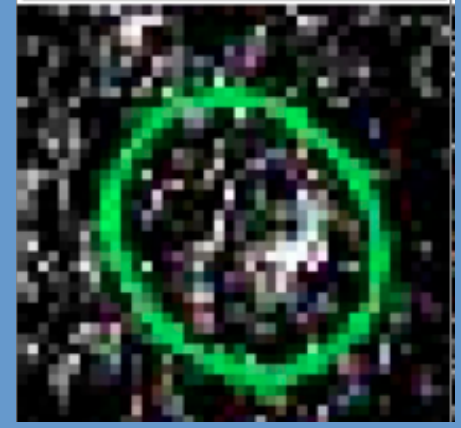
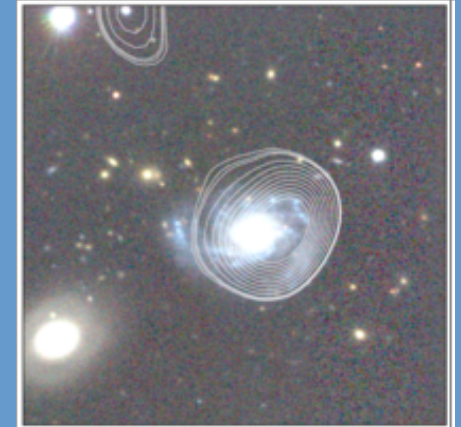
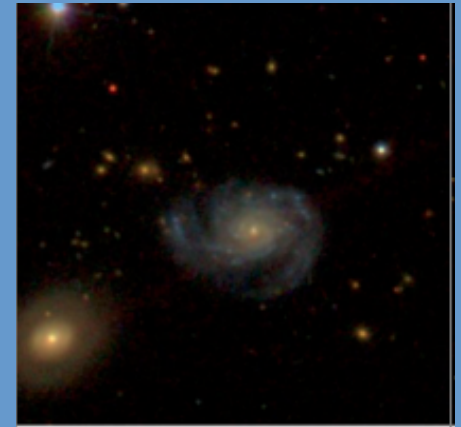
	No. of objects
DR2 detections	~280,000
DR2 extended sources	~30,000
DR2 extended sources in SDSS DR13 with >200 counts	~2,000
DR2 clusters in SDSS DR13 with >200 counts	~850
DR2 clusters in SDSS DR13 with >200 counts with good quality L_x, T_x	~570
DR2 clusters in SDSS DR13 with >200 counts with good quality z	~460

Optical confirmation



XMM point spread function (PSF) is strong function of off-axis angle \rightarrow Point sources (foreground AGN, stars) can appear as extended sources

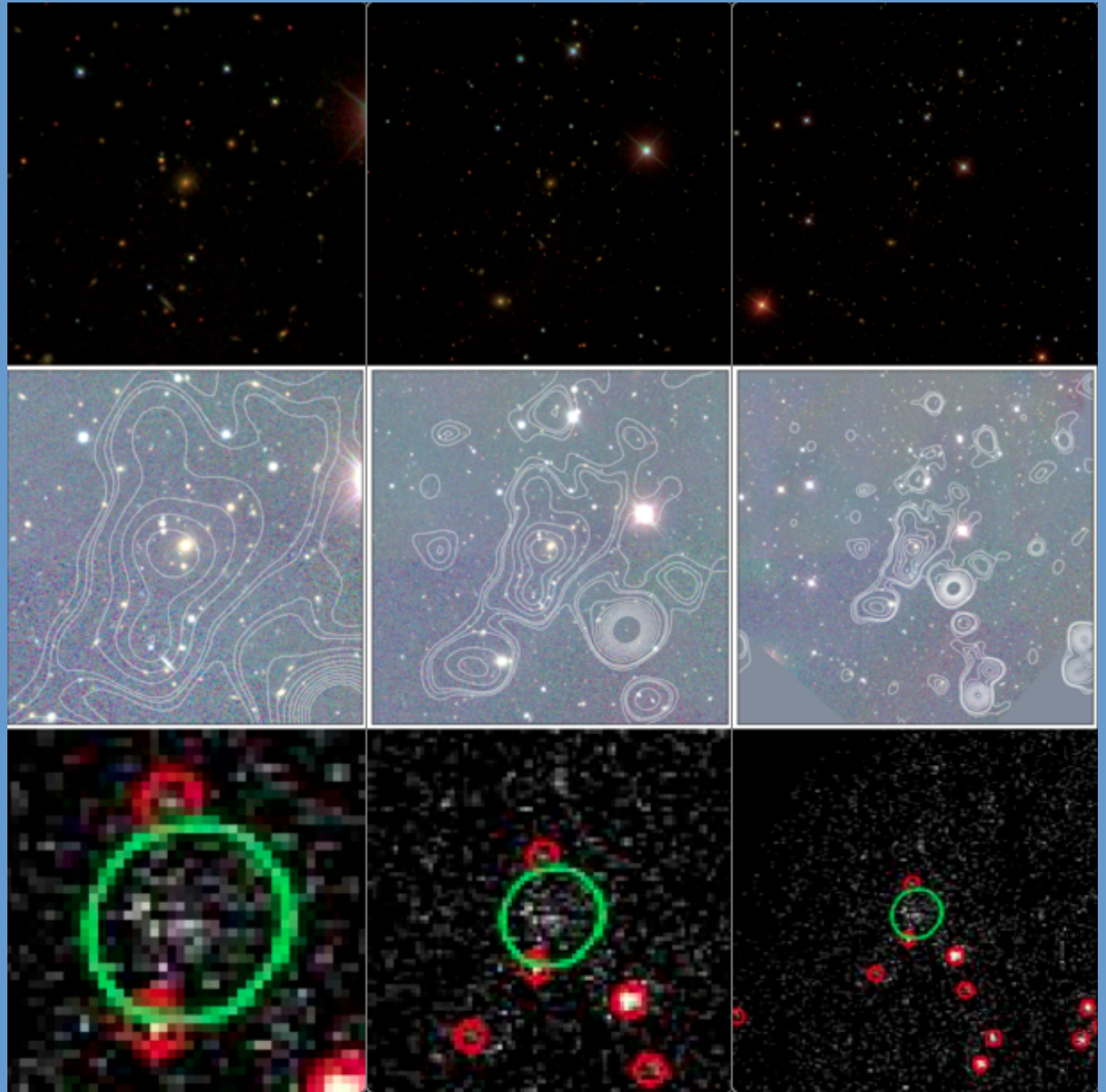
We need optical surveys to confirm our candidates \rightarrow SDSS is good for the northern hemisphere



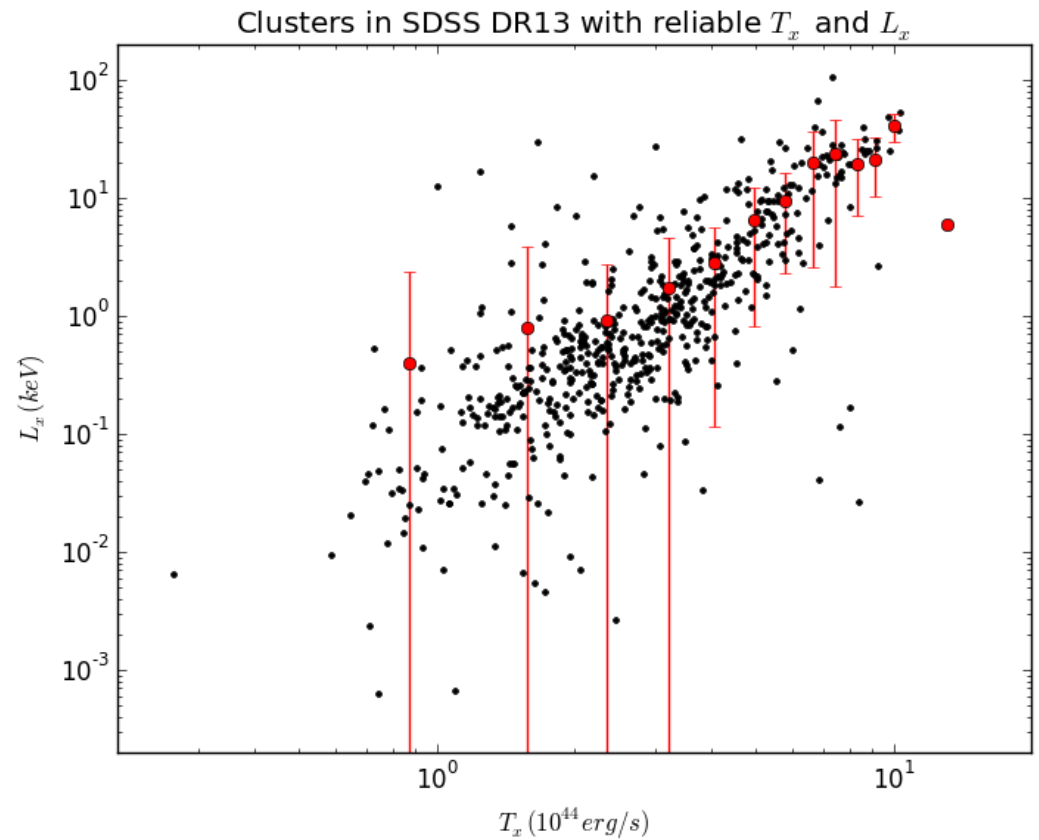
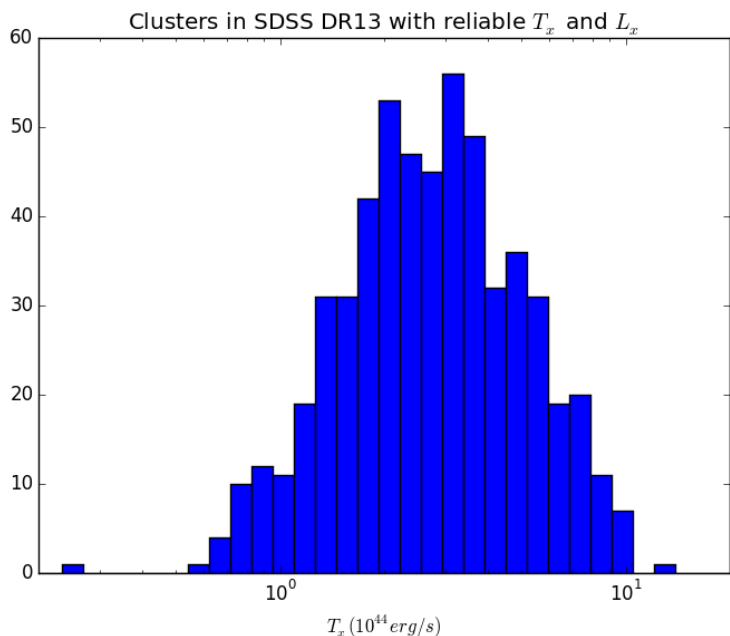
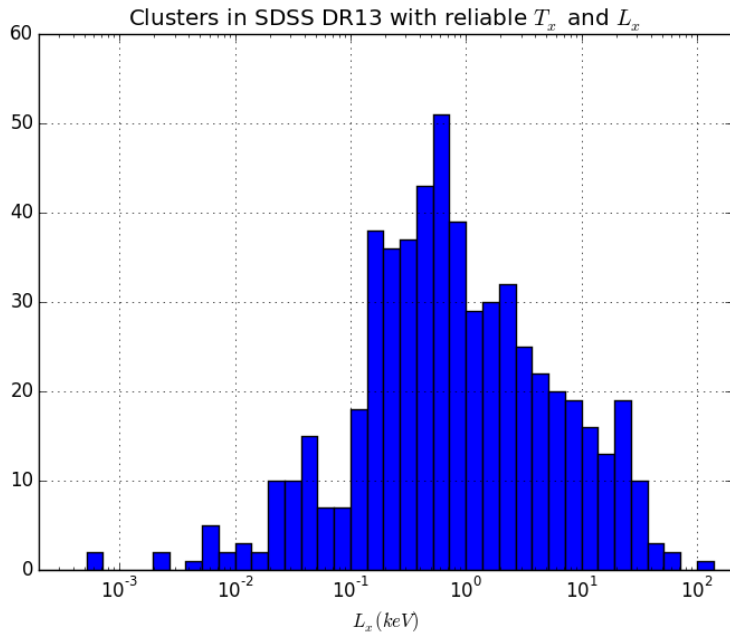
Optical confirmation

Cluster zoo
for the XCS
DR2 objects in
SDSS DR13

FOV: 3x3, 6x6,
12x12 arcsec

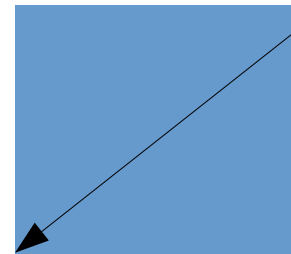
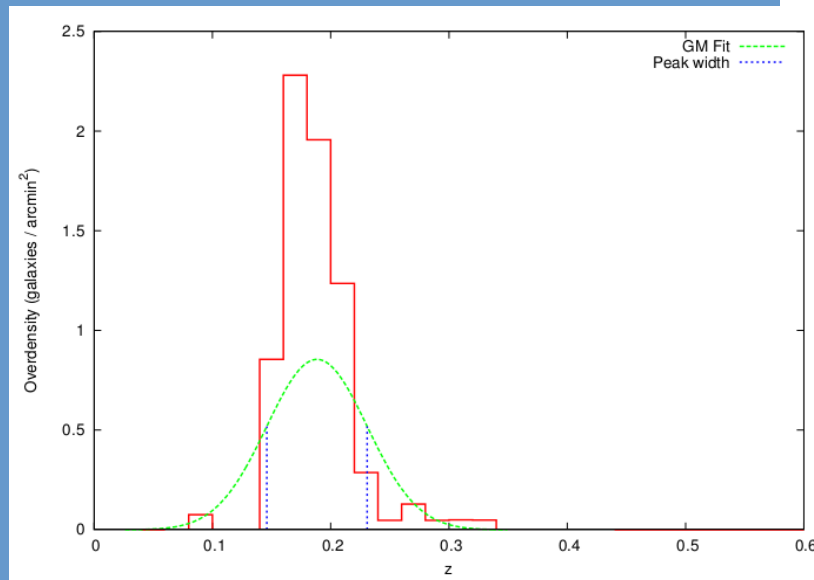
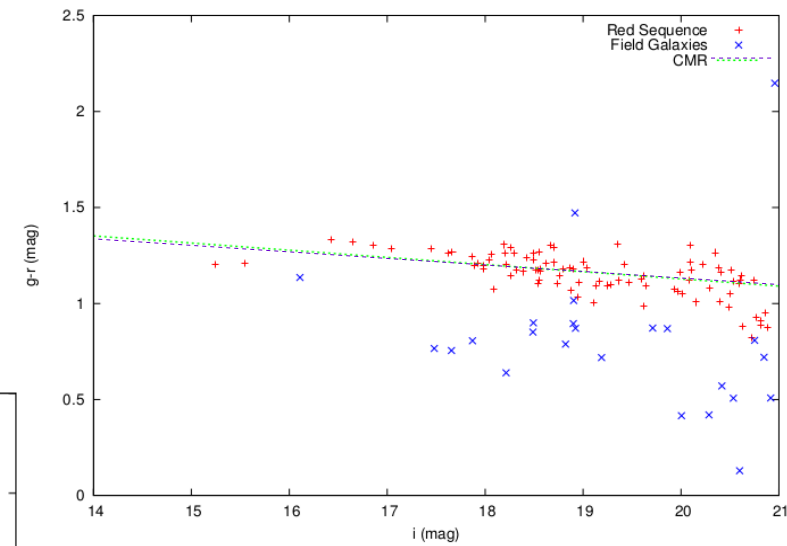
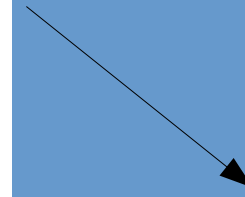
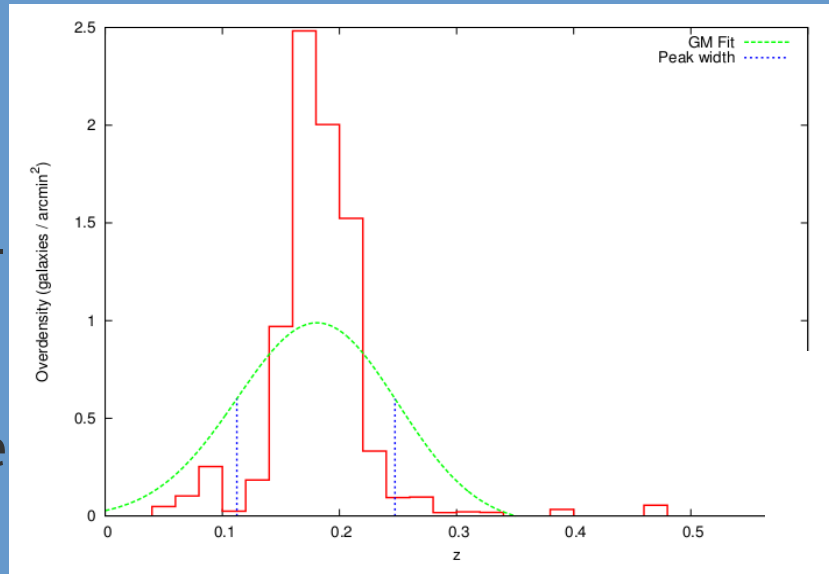


Optical confirmation



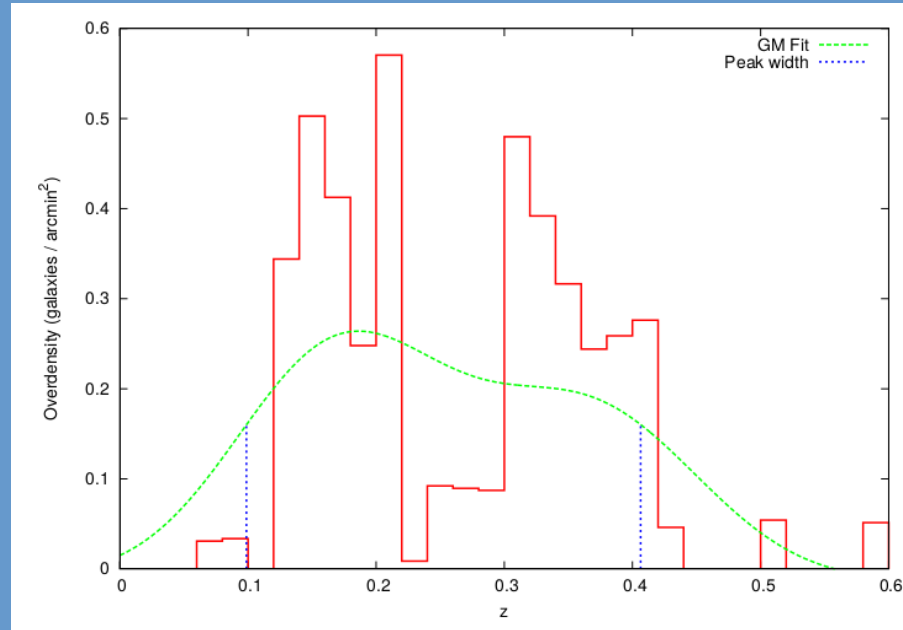
Cluster redshifts

- GMPhoRCC (Hood & Mann 2016)
- Two main cluster characteristics: red-sequence and BCG
- Traces the overdensities in the galaxy redshift distribution
- Input: galaxy survey and cluster coordinates
- Output: optical cluster redshift, colour-magnitude relation, cluster richness
- Quality flags

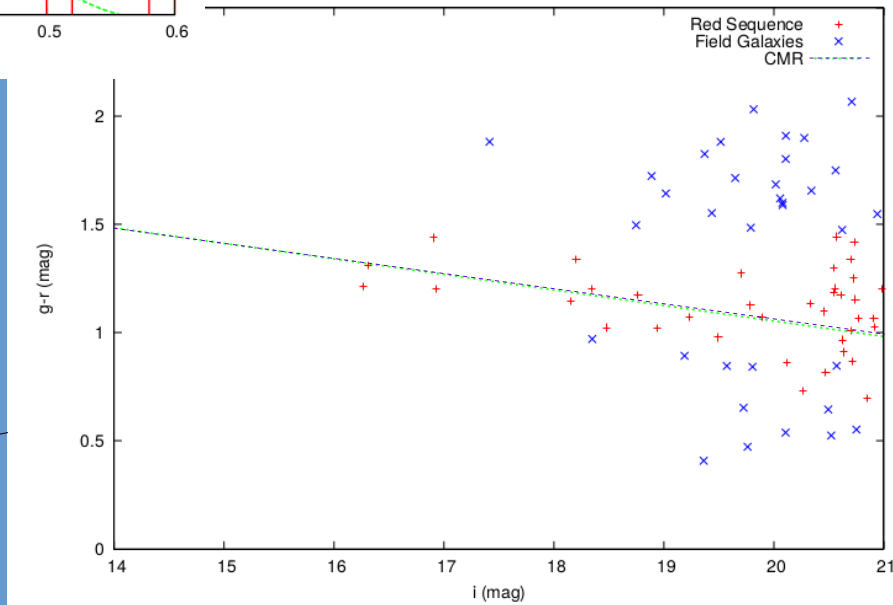
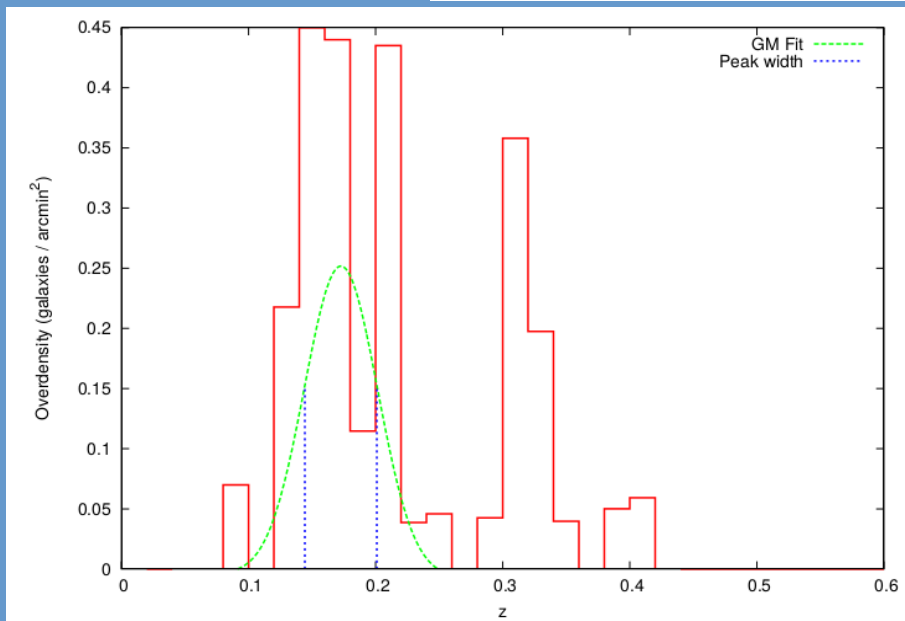


Cluster redshifts

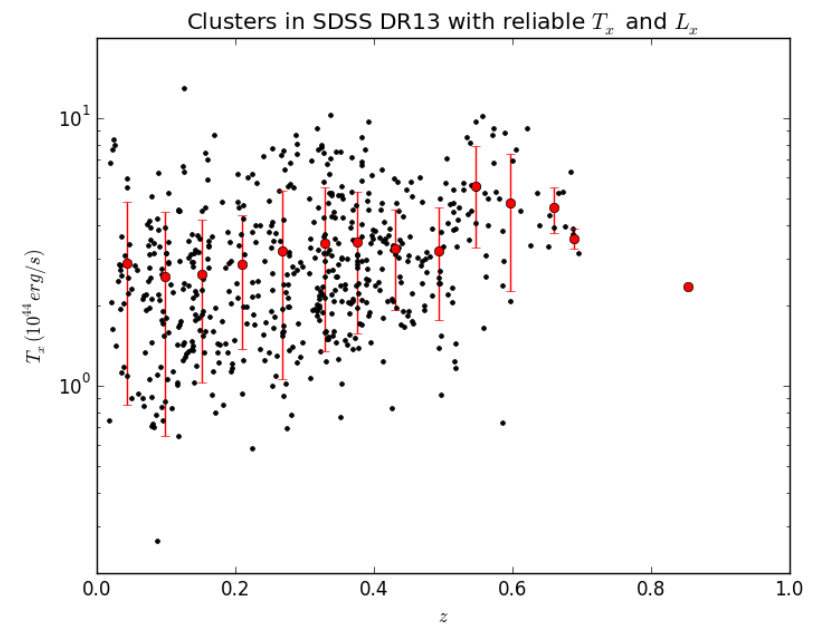
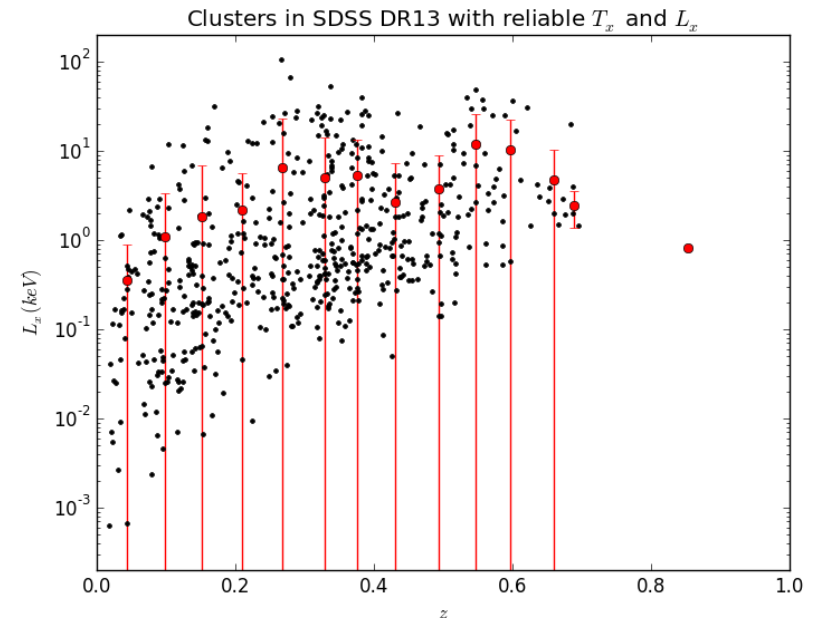
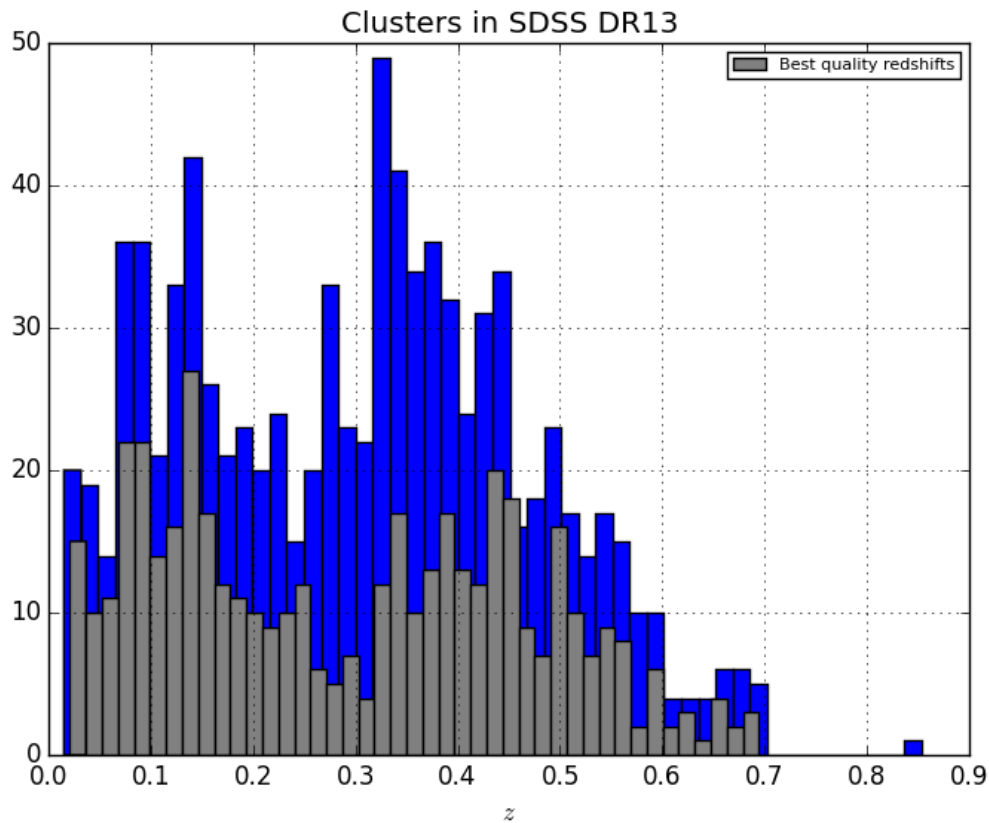
Analyses
several
candidates



$$z_{rs} = 0.172$$
$$z_{spec} = 0.179$$

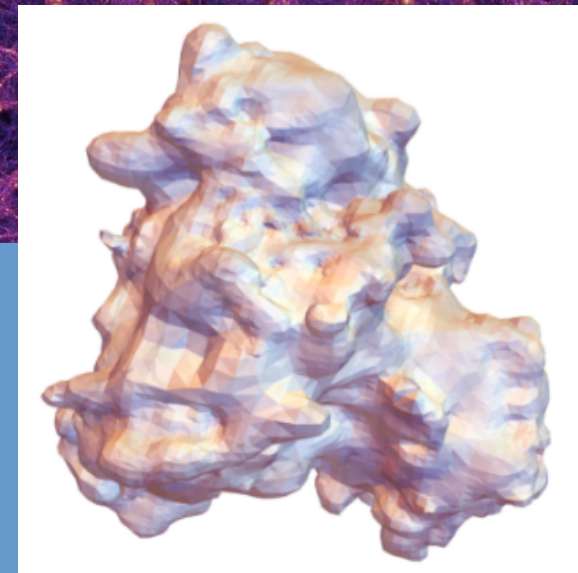
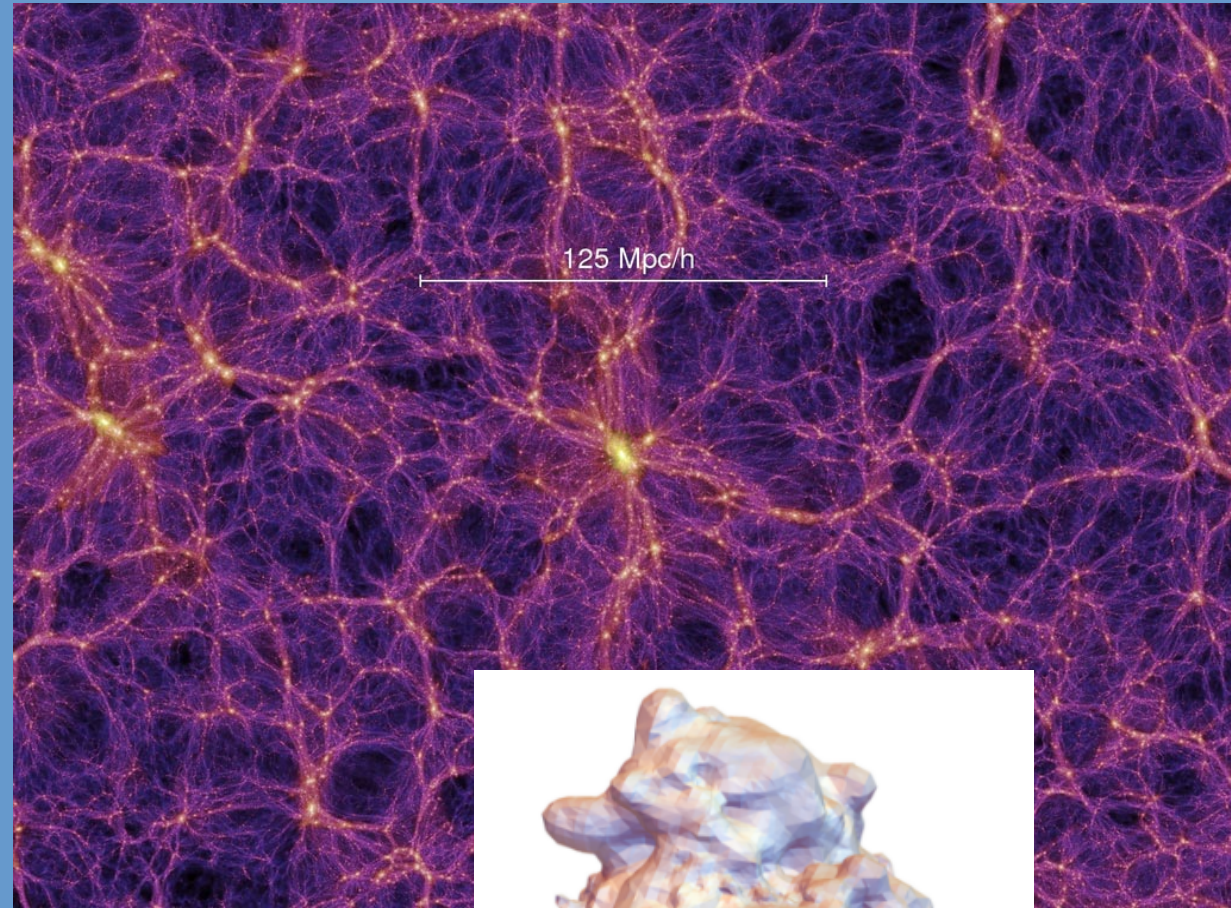


Cluster redshifts



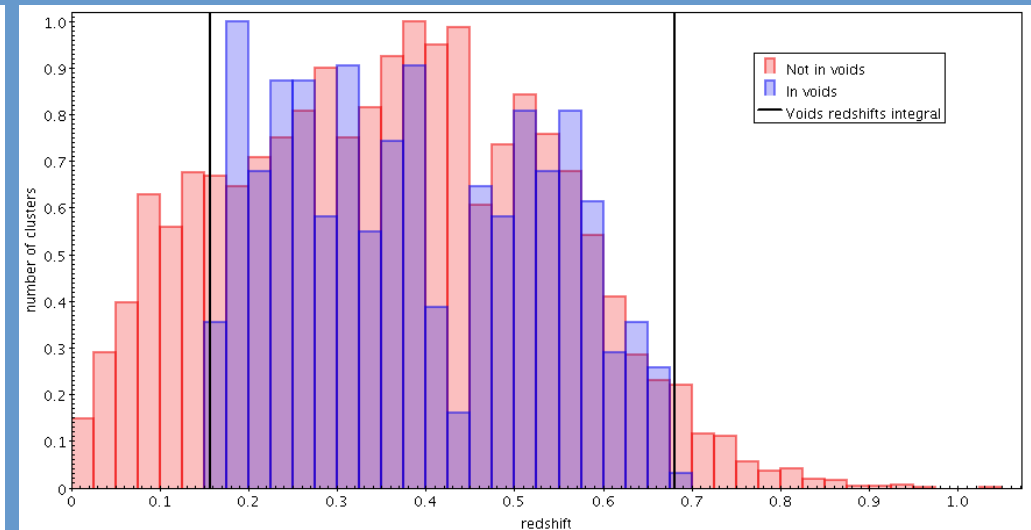
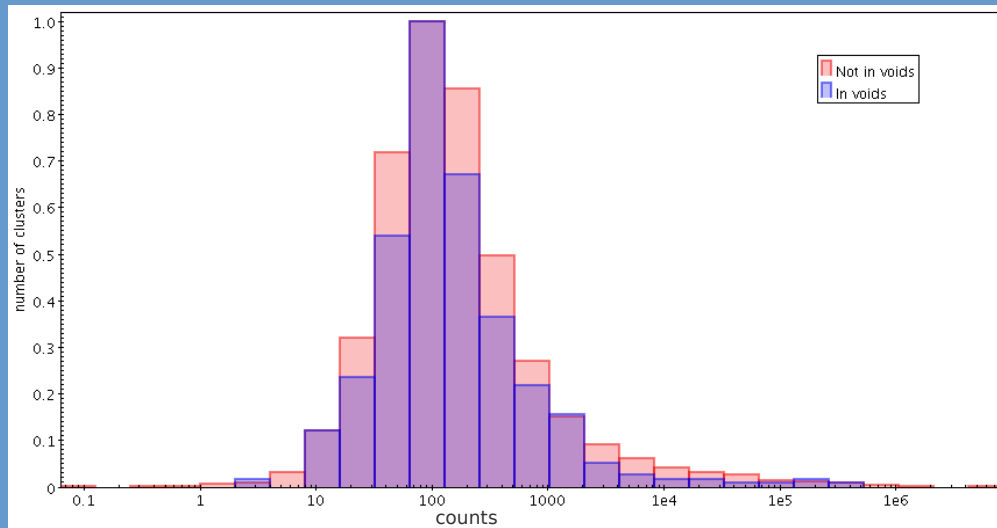
Clusters in voids

- Underdense regions in LSS, complicated 3D void shapes
- Void catalogue in BOSS DR12 (Nadathur 2016)
- Different merger histories clusters in voids and not \rightarrow developed different properties
- Expect less massive, less luminous clusters inside voids, more spread temperatures



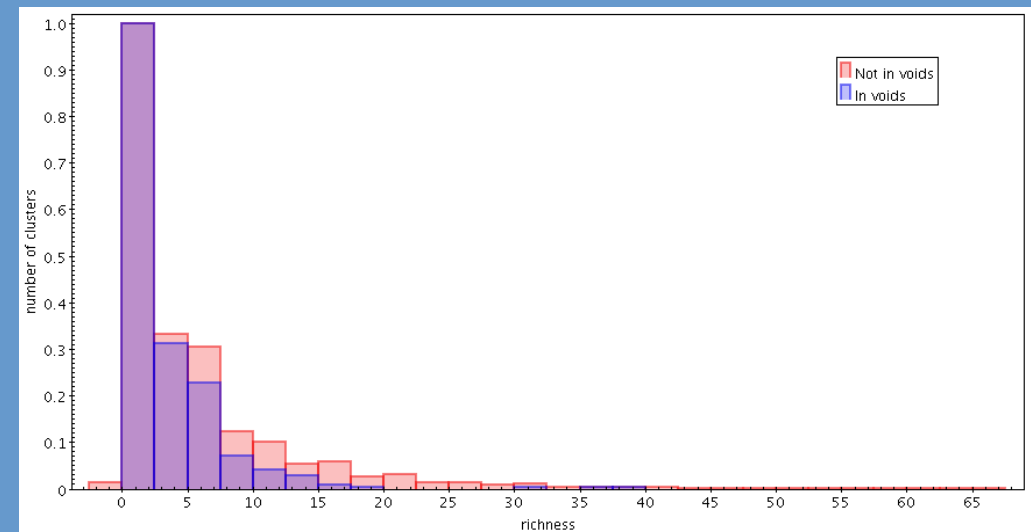
Clusters in voids

First comparisons of clusters inside and outside voids



Developed algorithm which identifies if a cluster lies inside a void or not

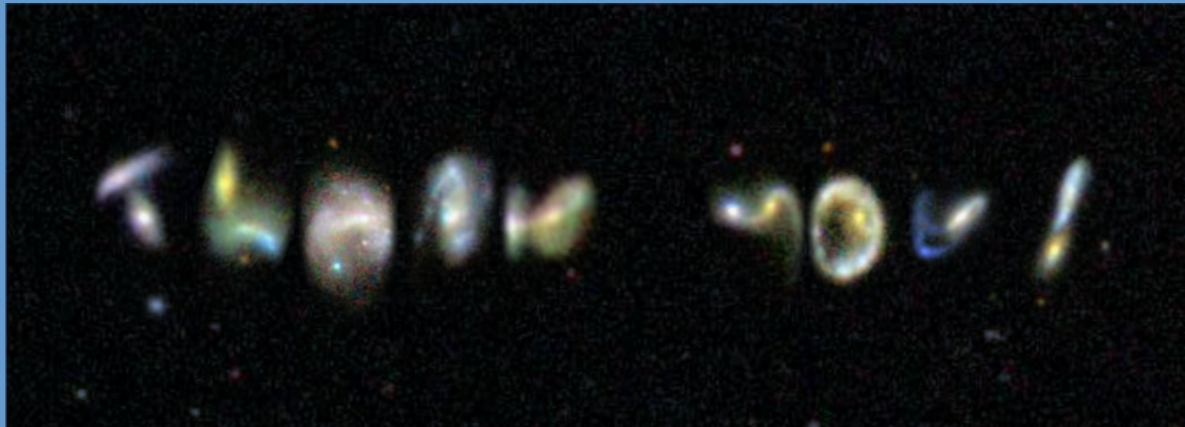
Assumed spherical voids, no counts limit



Future work

- Publish DR2
- Introduce 3D void shape
- Compare with Millennium simulation voids
- Introduce different flux limits in our cluster sample and study the effect of the resulting selection bias

XCS DR2 is coming soon
Stay tuned!



Extra material

Red sequence band	redshift range
$g - r$	$0.00 \leq z < 0.45$
$r - i$	$0.35 \leq z < 0.75$
$i - z$	$0.65 \leq z$

