The background of the slide is a dark, reddish-brown field filled with numerous small, bright yellow and white stars. A prominent, glowing purple filamentary structure, resembling a galaxy or a nebula, stretches diagonally across the center of the image. The text is overlaid on this background in a bright cyan color.

Not another TLA!!
ERQs from SDSS-III BOSS
and WISE; the WW4C and
JWST MIRI ERS

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MJD 57576

*Galaxies have two key energy sources:
nuclear fusion and energy liberated in a
strong gravitational field.*

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poorly understood component in galaxies.**

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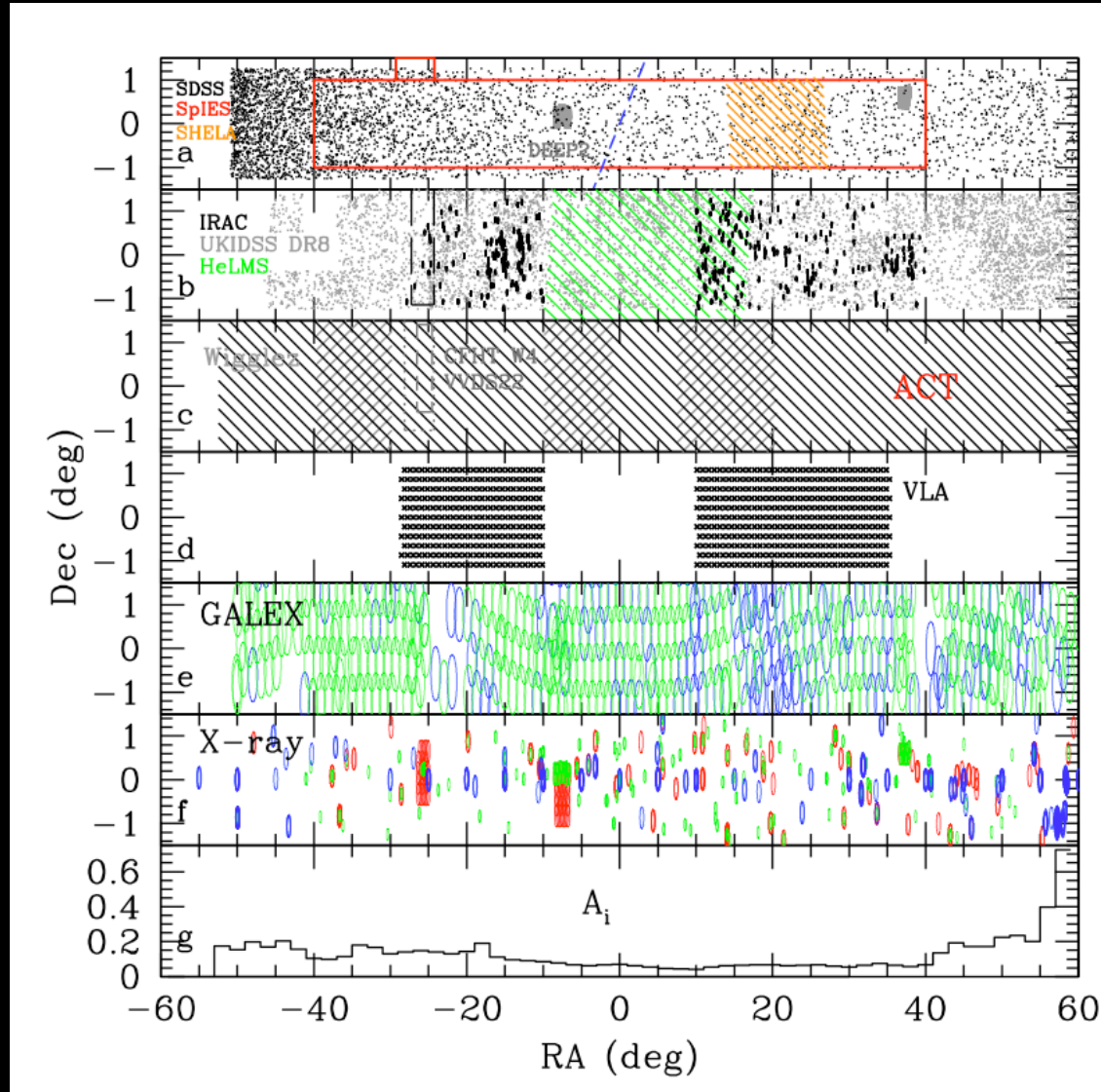
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**Key Question: At Cosmic Noon*, what is the
relation between nuclear fusion and gravitation
accretion, and how does it influence the galaxy
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***Which comes shortly after Cosmic Elevenses ;-)**

SpIES: Spitzer-IRAC Equatorial Survey

SDSS Stripe 82 field



Science Case:
Quasar LF +
Clustering at
 $z \sim 3.5$ and 4.5

Obscured
AGN

$z > 6$ QSOs

Massive red
galaxies at $z > 1$

Cool stars

Spitzer
Space
Telescope

Exploration
Science

Programme
820hrs awarded

Catalogs and
Survey Paper:

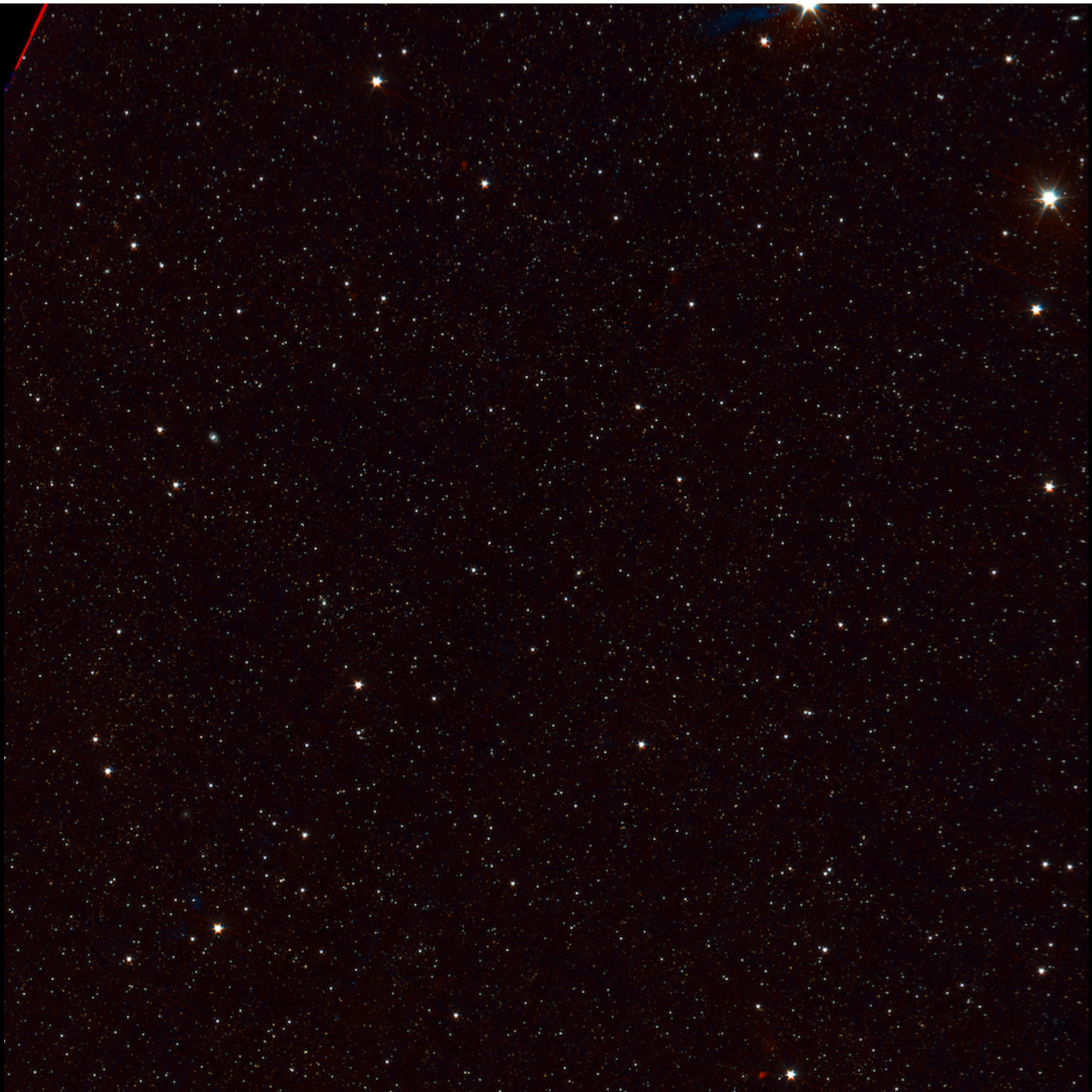
Timlin, NPR et al.
2016, ApJS,
225, 1

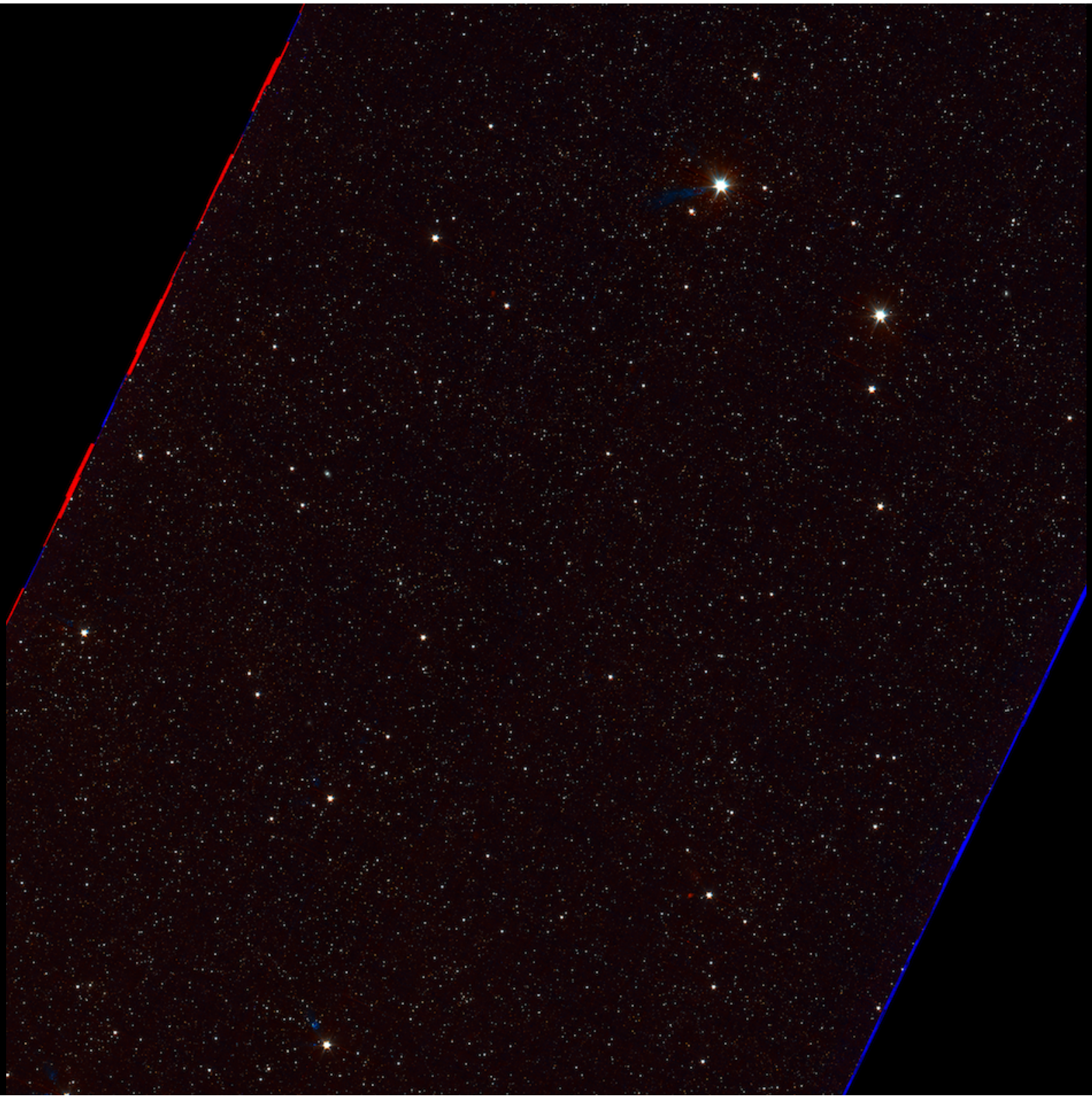
110 deg² 3.4 μ m (5 μ Jy) and 4.5 μ m (7 μ Jy)

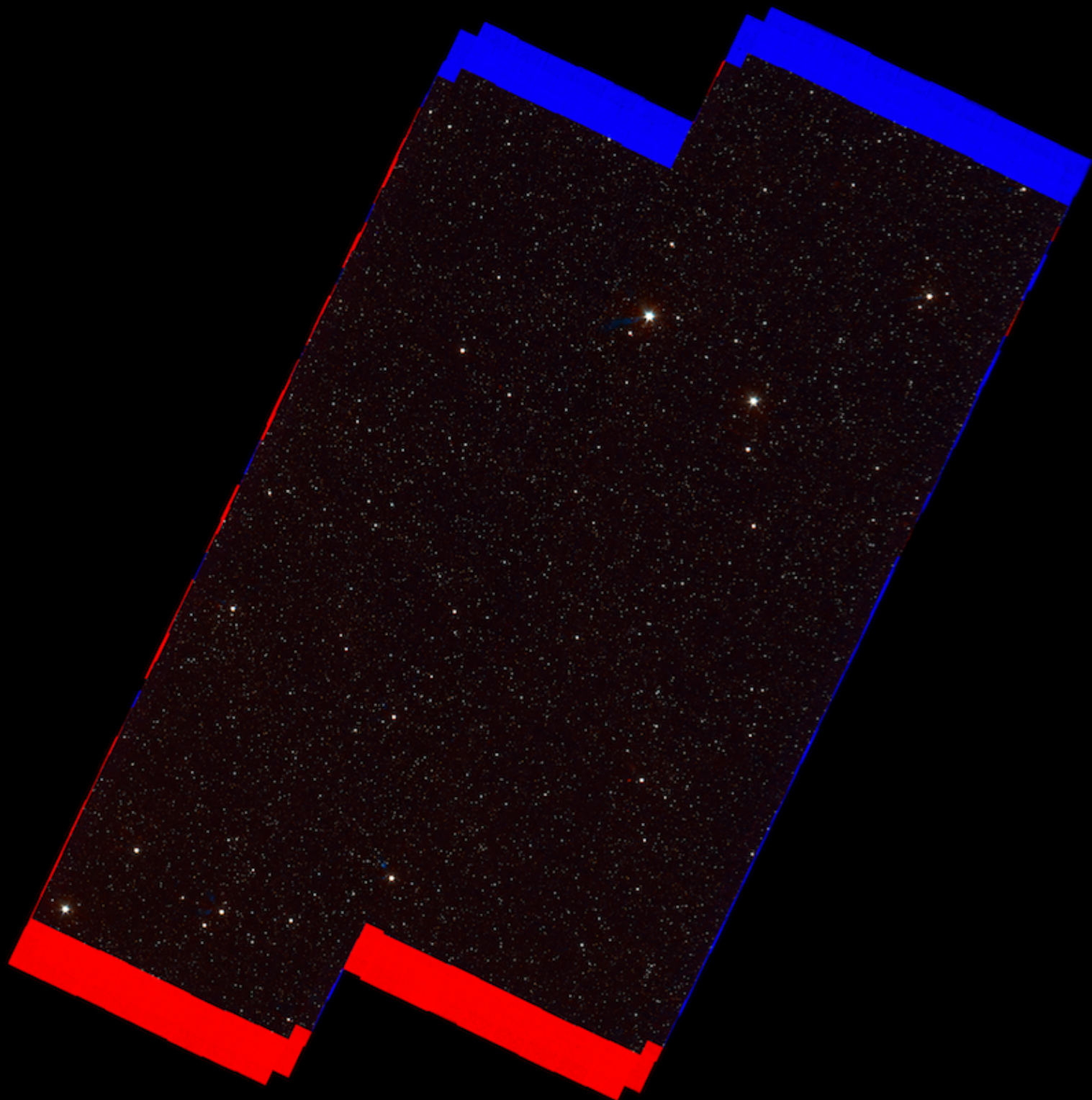


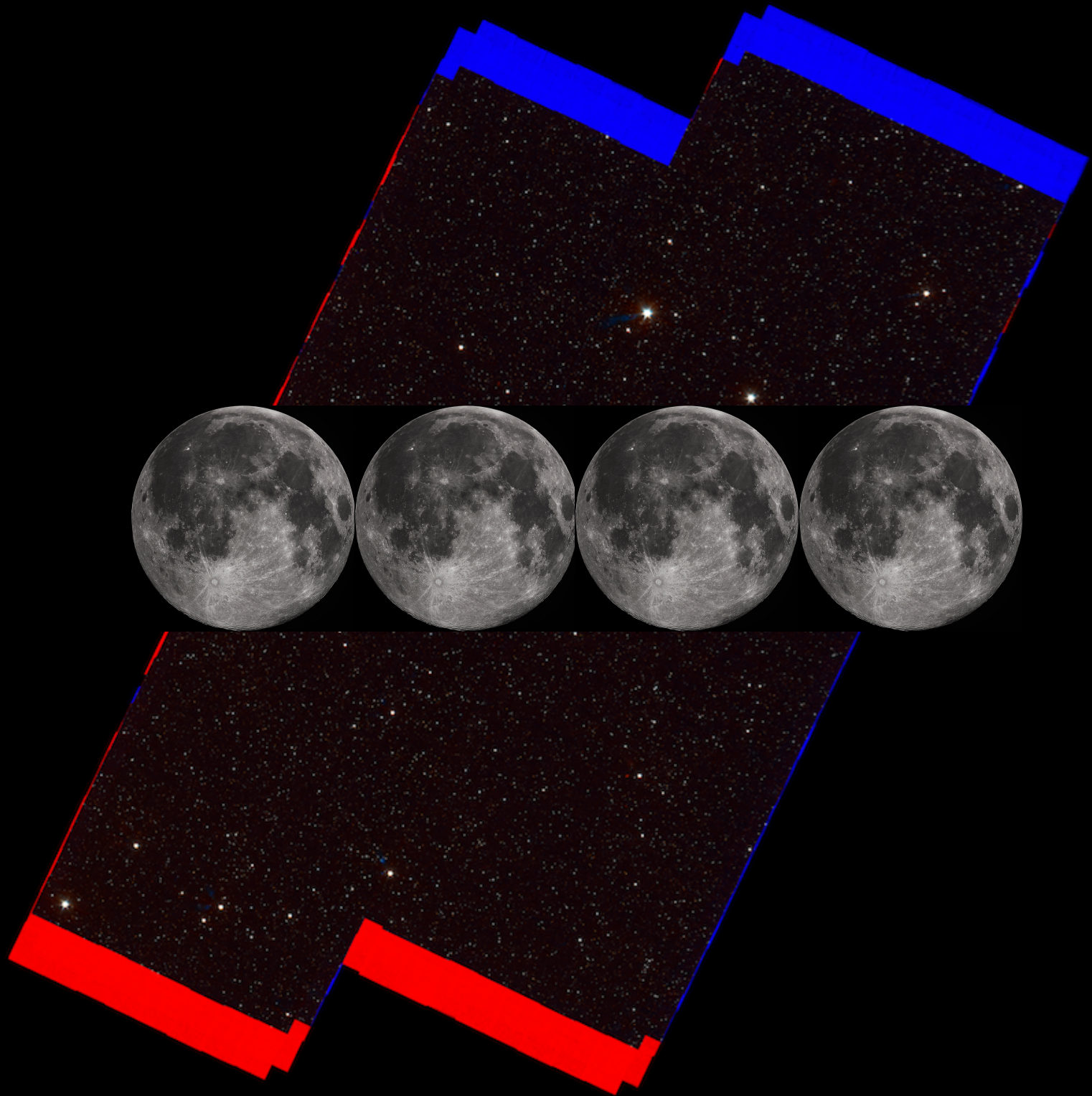








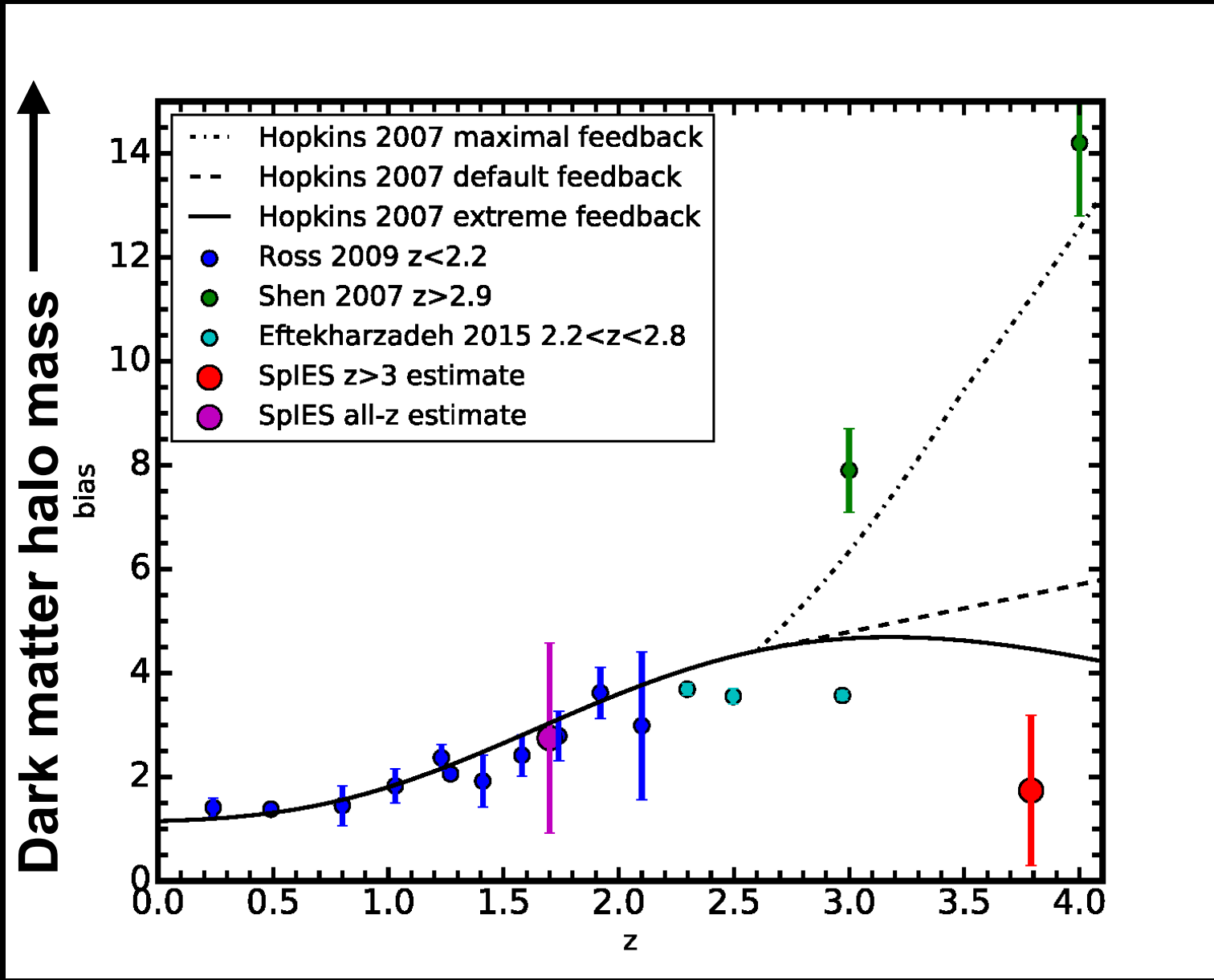




$\sim 1/30^{\text{th}}$
full SpIES

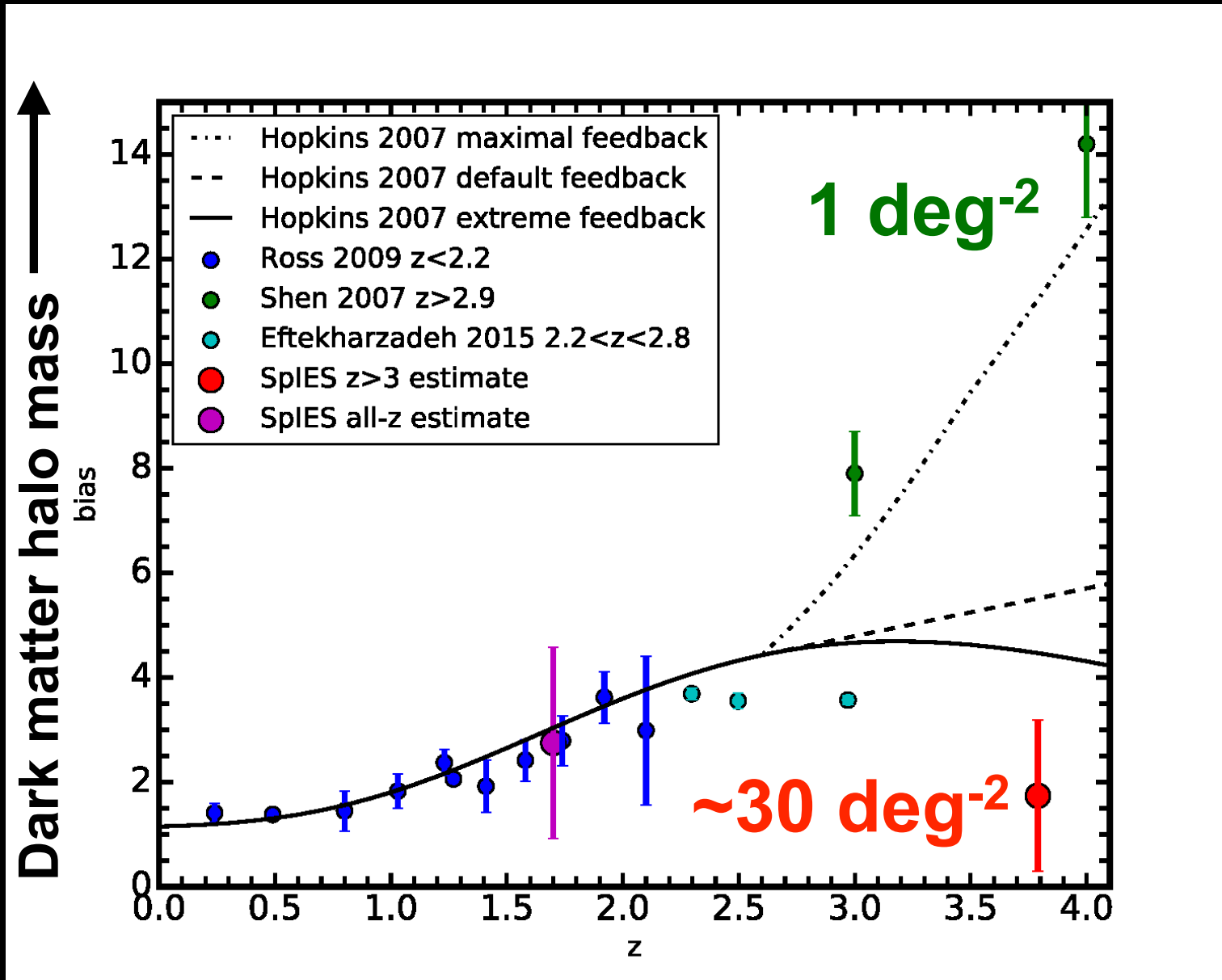


SpIES: First IR QSO Clustering at $z > 2$



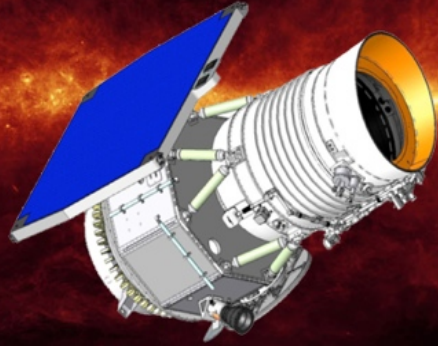
Timlin, *NPR et al.*, 2016b, *ApJ*, in prep.

SpIES: First IR QSO Clustering at $z > 2$



Timlin, *NPR et al.*, 2016b, *ApJ*, in prep.

Wide Infrared Survey Explorer (WISE)



3.4, 4.6, 12, 22 μm
0.08, 0.11, 1, 6 mJy
All-Sky Survey



Wright et al. (2010)
wise.ssl.berkeley.edu

Extremely Red Quasars

65 / 255,950 (0.025%) of
SDSS DR7 + BOSS DR10 have $r-[22] > 14.0$

2'

$z=0.84$

138.75714, 24.30336

SDSS *gri*

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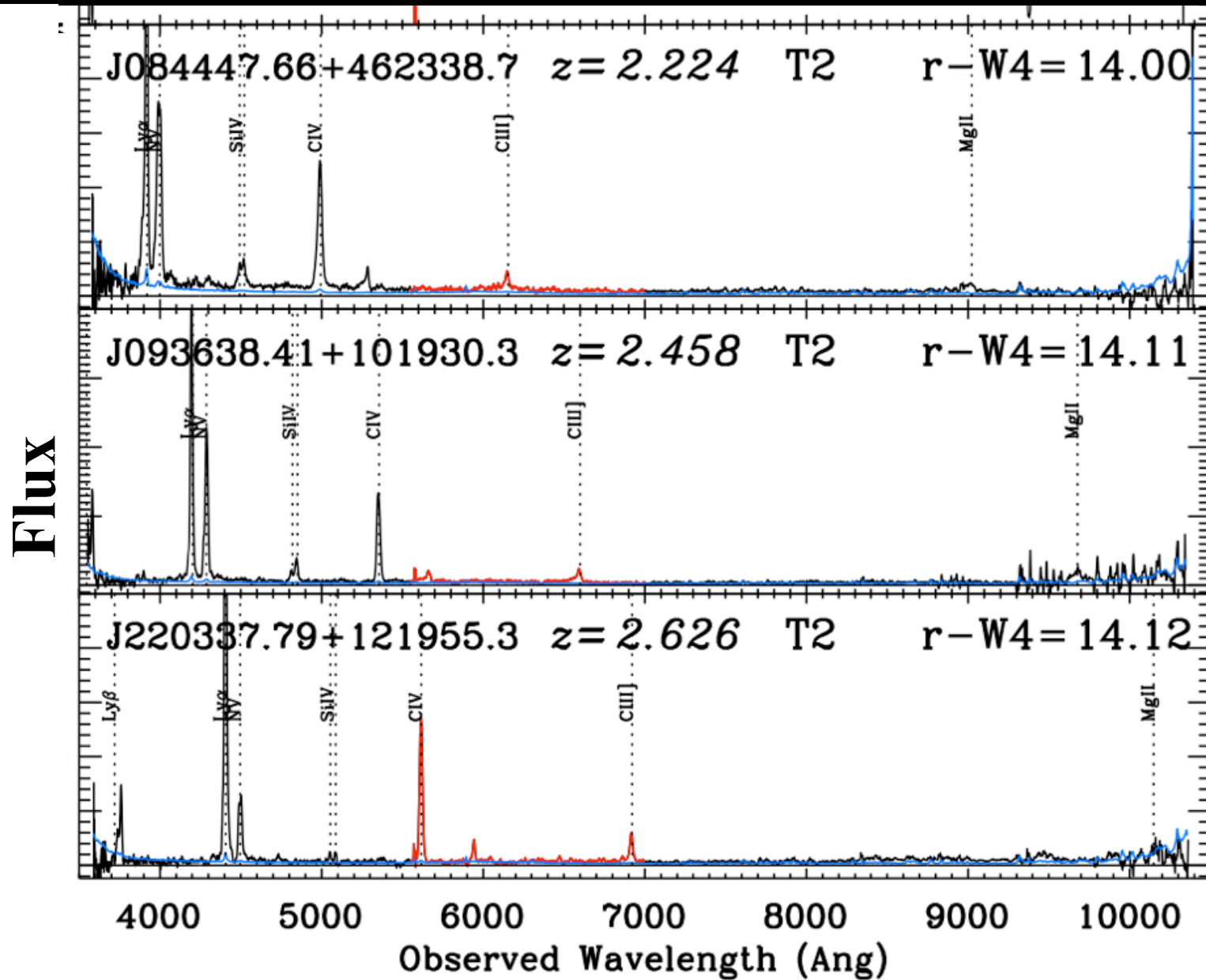
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SDSS *gri*

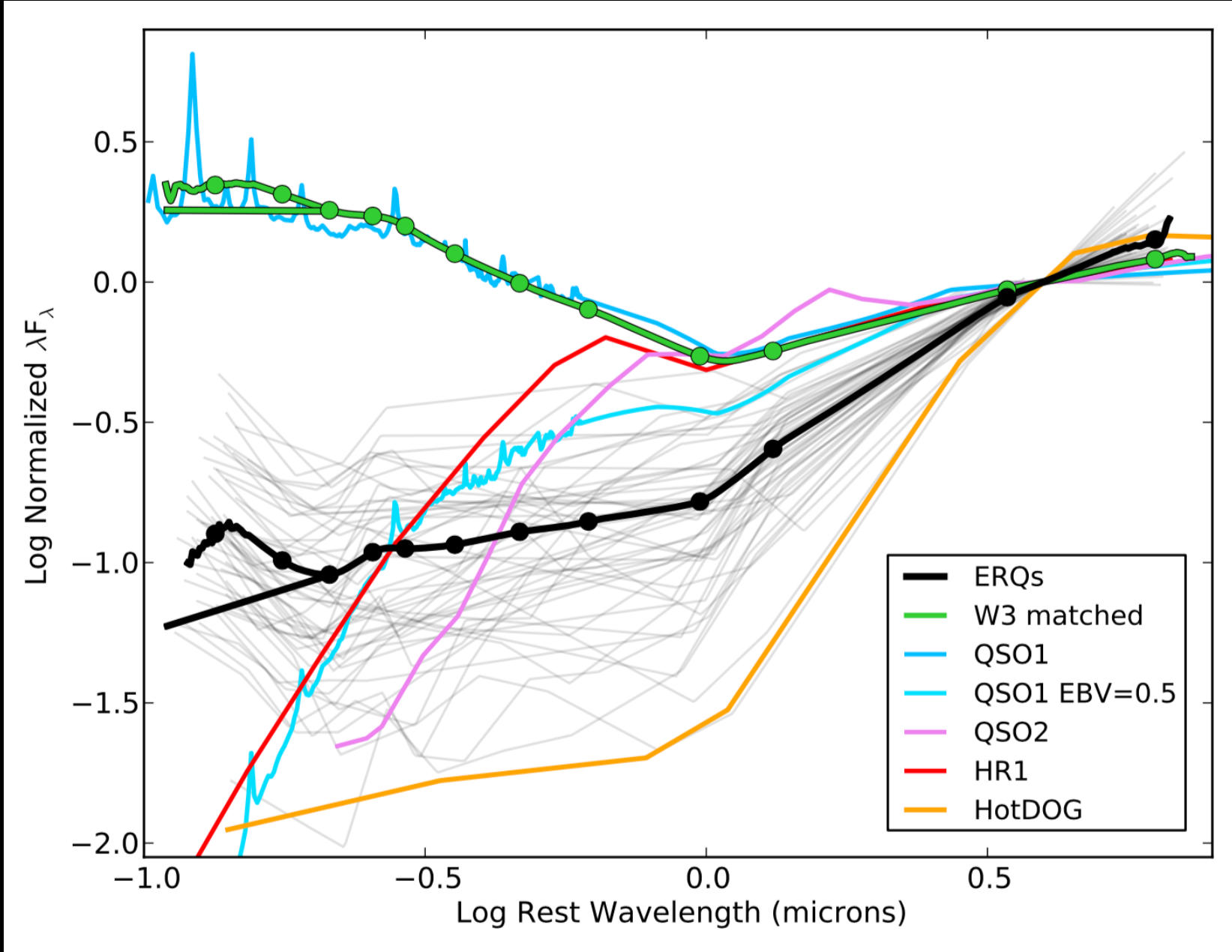
$r-[22] = 15.5$

WISE W1W2W3

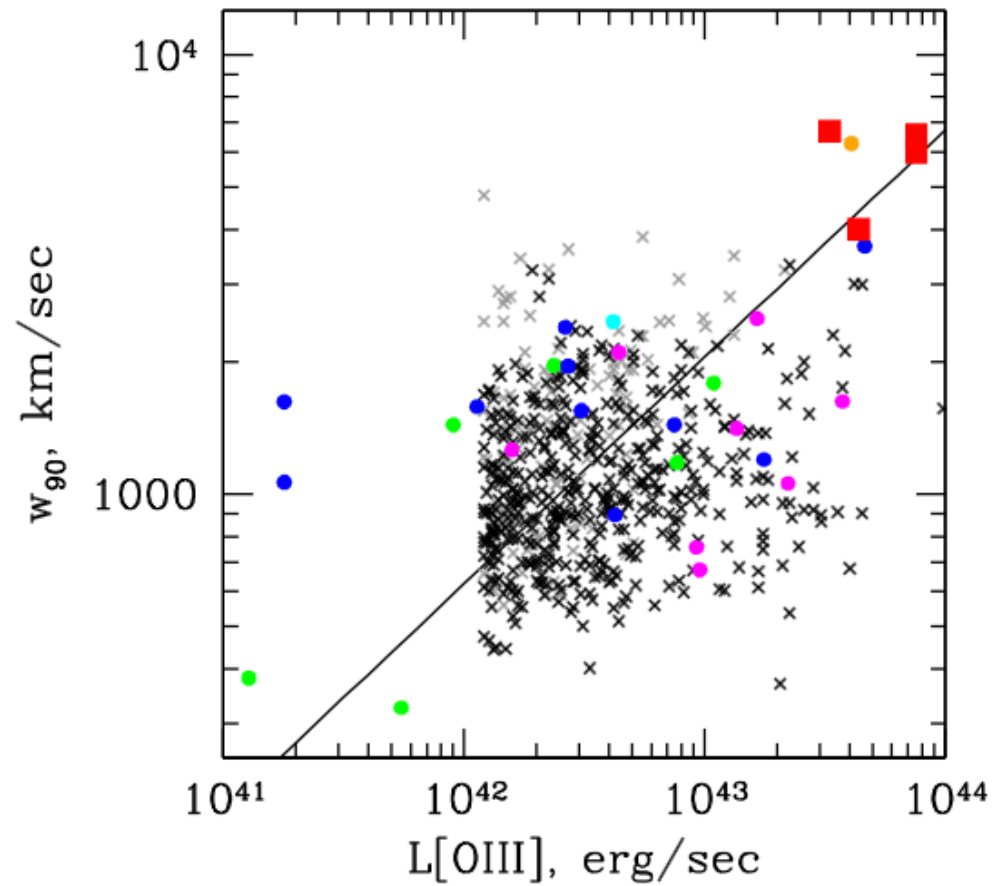
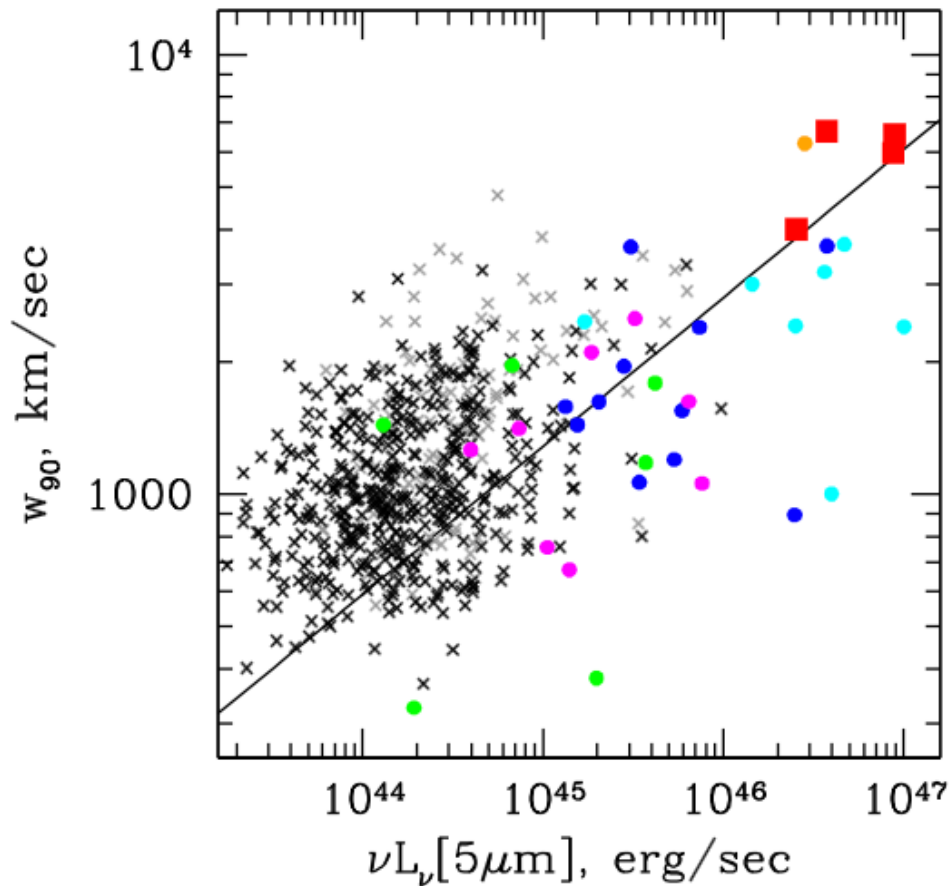
"Type 2" QSOs at high- z



ERQ SEDs: An Evolutionary trend?



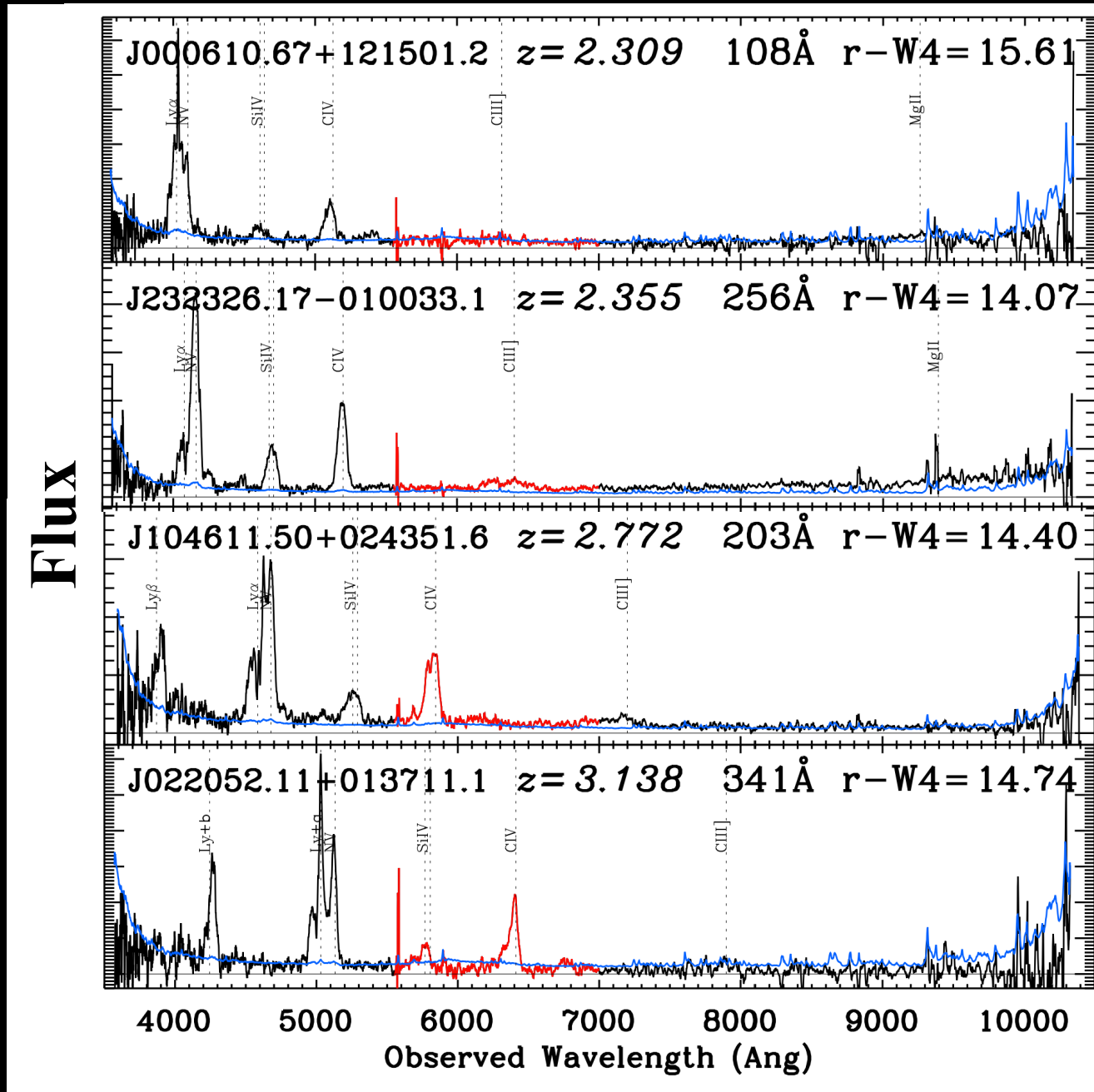
Emission Lines of ERQs: Feedback in Action



$\nu L_{\nu}[5\mu\text{m}]$, erg/sec
 10^{44} 10^{45} 10^{46} 10^{47}

$L[\text{OIII}]$, erg/sec
 10^{41} 10^{42} 10^{43} 10^{44}

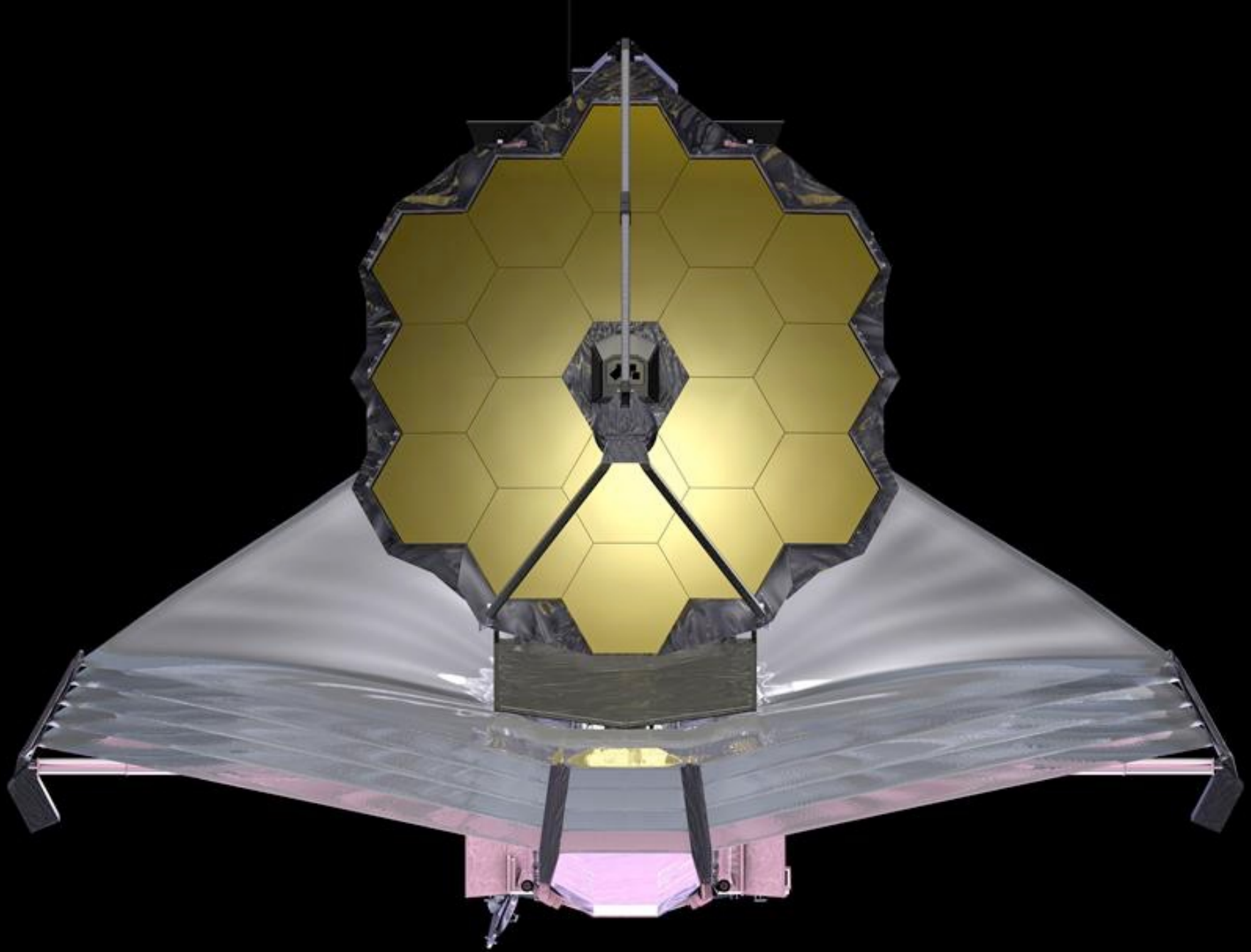
Very Unusual Spectra from ERQs....



The WISE W4 Compendium

- ✓ All 40 million detected objects in the WISE W4-band (20-28um); concentrating on objects with dust emission and AGN SEDs
- ✓ $z \sim 2.5$ object detected in WISE W4 has $L_{IR} \gtrsim 10^{13.5} L_{\odot}$
- ✓ *(Might) Become a target catalogue for JWST MIRI spectroscopy...*

WISE band combination	det_bit	Number of objects	Percentage of AllWISE
W1-W2-W3-W4	15	25 882 083	3.5
W1-W2-W4	11	11 309 923	1.5
W1-W4	9	2 347 472	0.3
W1-W3-W4	13	859 426	0.1
W3-W4	12	454 160	0.1
W4	8	35 818	<0.1
W2-W3-W4	14	35 528	<0.1
W2-W4	10	15 556	<0.1
W4-any		40 939 966	5.5



Coming Summer 2019

