



Galactic Clusters Survey: Update

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(on behalf of the GCS WG)



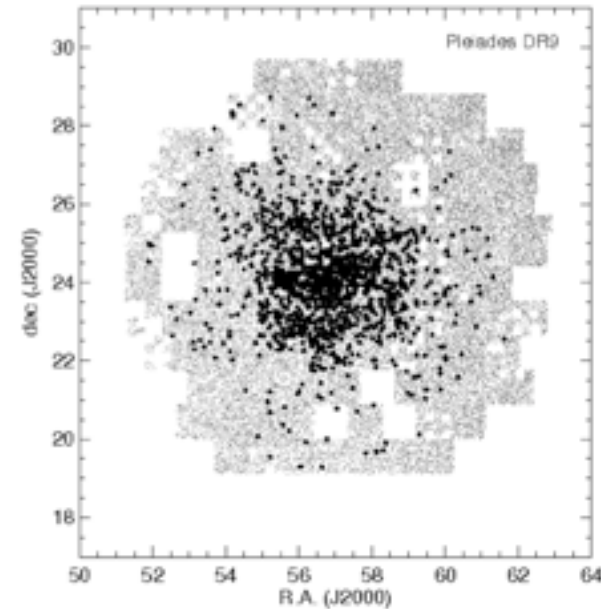
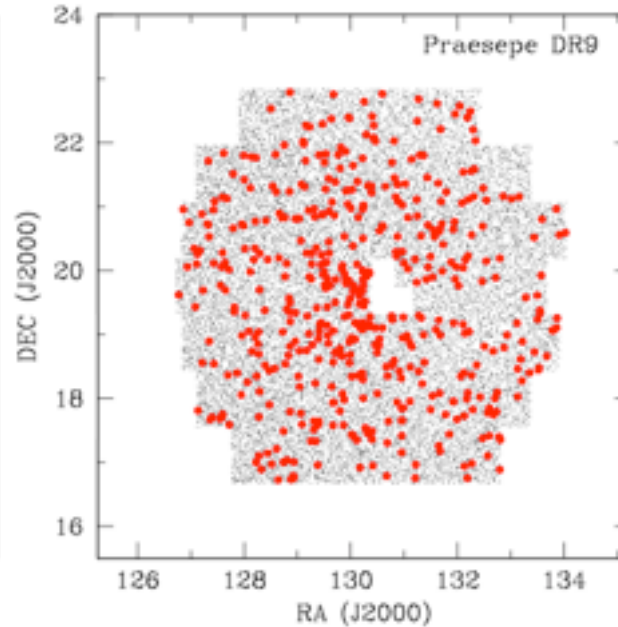
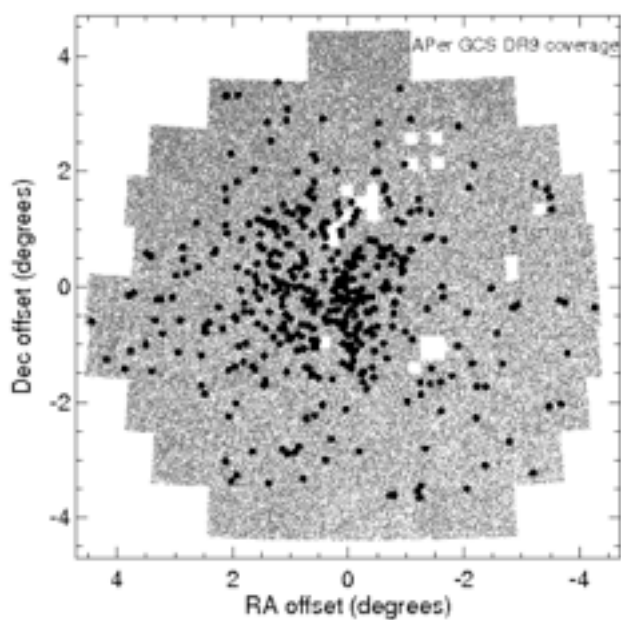
Survey coverage (%)

Target	Area deg ²	Z	Y	J	H	K×2
IC-4665	3	100	100	100	100	100
Pleiades	79	95	94	96	89	92
AlphaPer	50	100	100	100	100	100
Praesepe	28	83	82	80	82	82
Tau-Auriga	218	100	100	100	43	100
Orion	154	17	18	18	53	*53
Upper Sco	154	35	34	34	100	100
Perseus OB2	13	100	100	100	100	100
Coma Ber	79	92	91	96	100	100
Hyades	291	7	6	6	4	*100

NB:

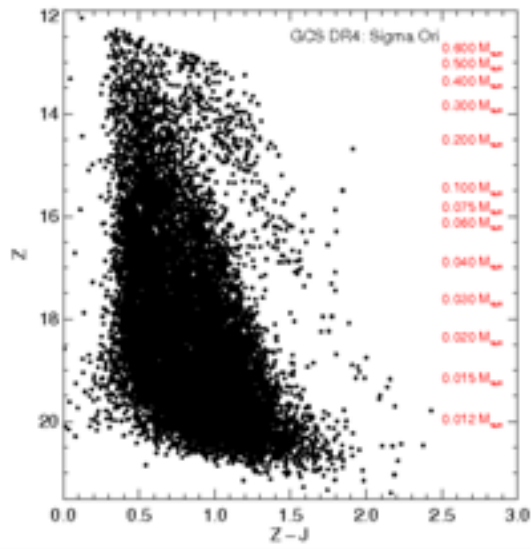
- includes processed data to end Jan 2012, but QC only to end Jan 2011 (for DR9)
- No K2 in Orion and Hyades suspended
- 77% complete (compared to original plan) as at April 2012.

Coverage (cont)

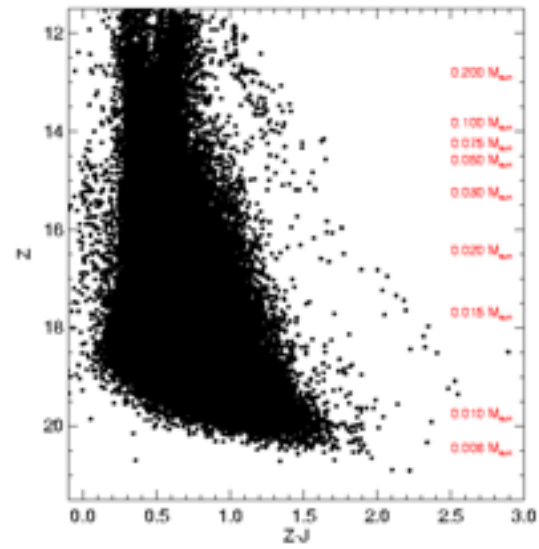


- Good DR9 coverage (as at end Semester 10B)
- Some gaps being filled (up to 12A)
- More extensive coverage than legacy optical photographic surveys
- Much larger samples than deep pencil-beam surveys

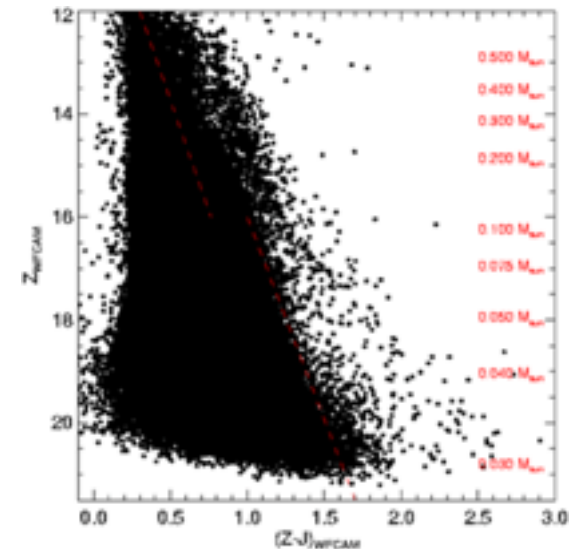
Cluster VLMS/BD sequences



Ages: ≥ 1 Myr

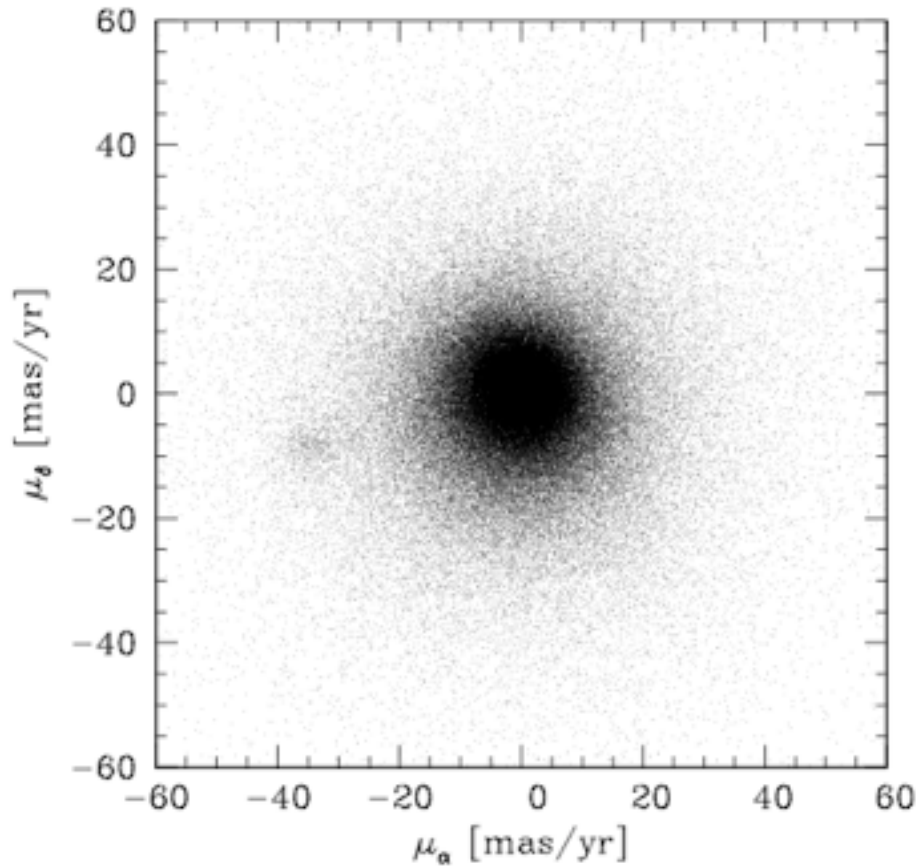


~ 10 Myr



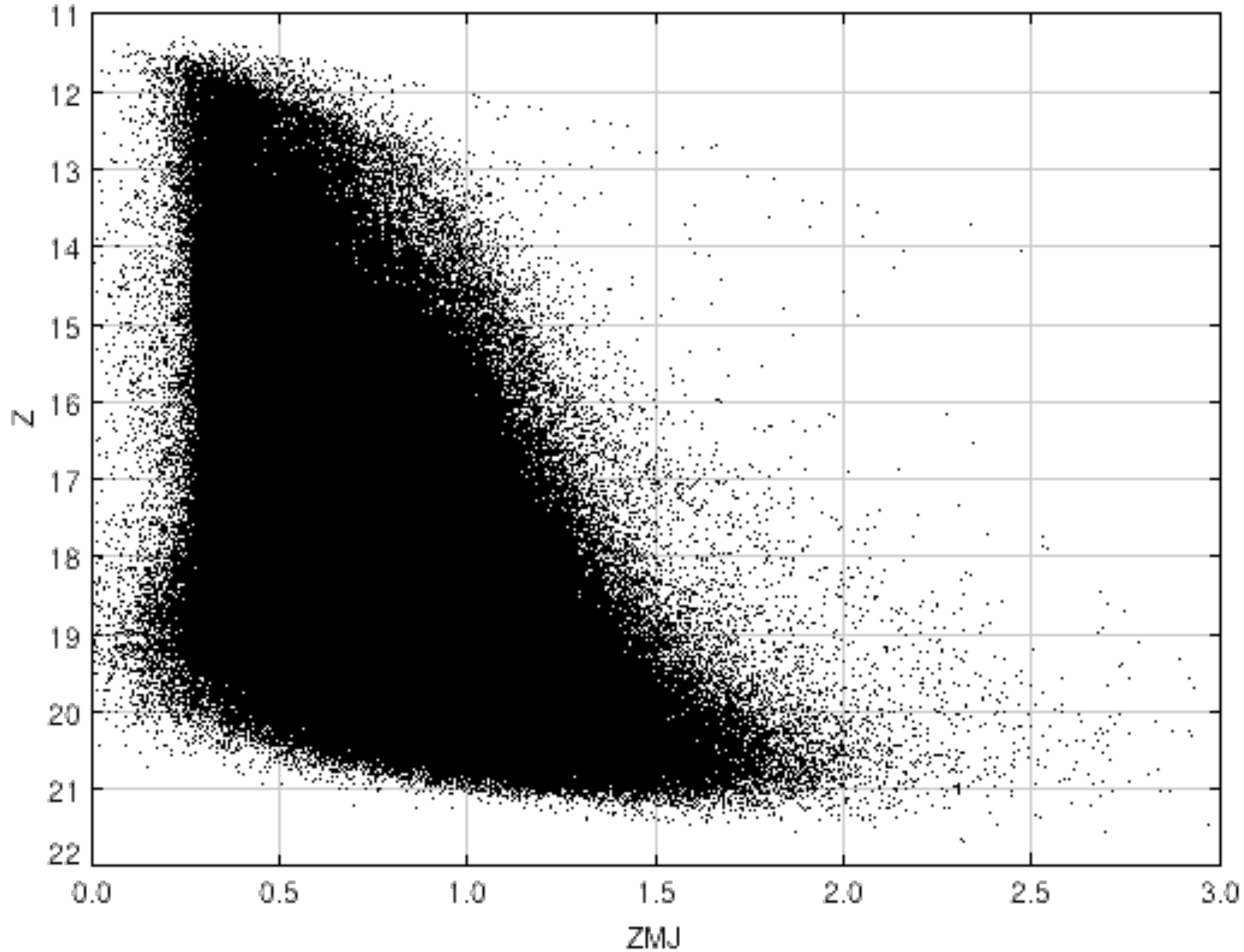
~ 100 Myr

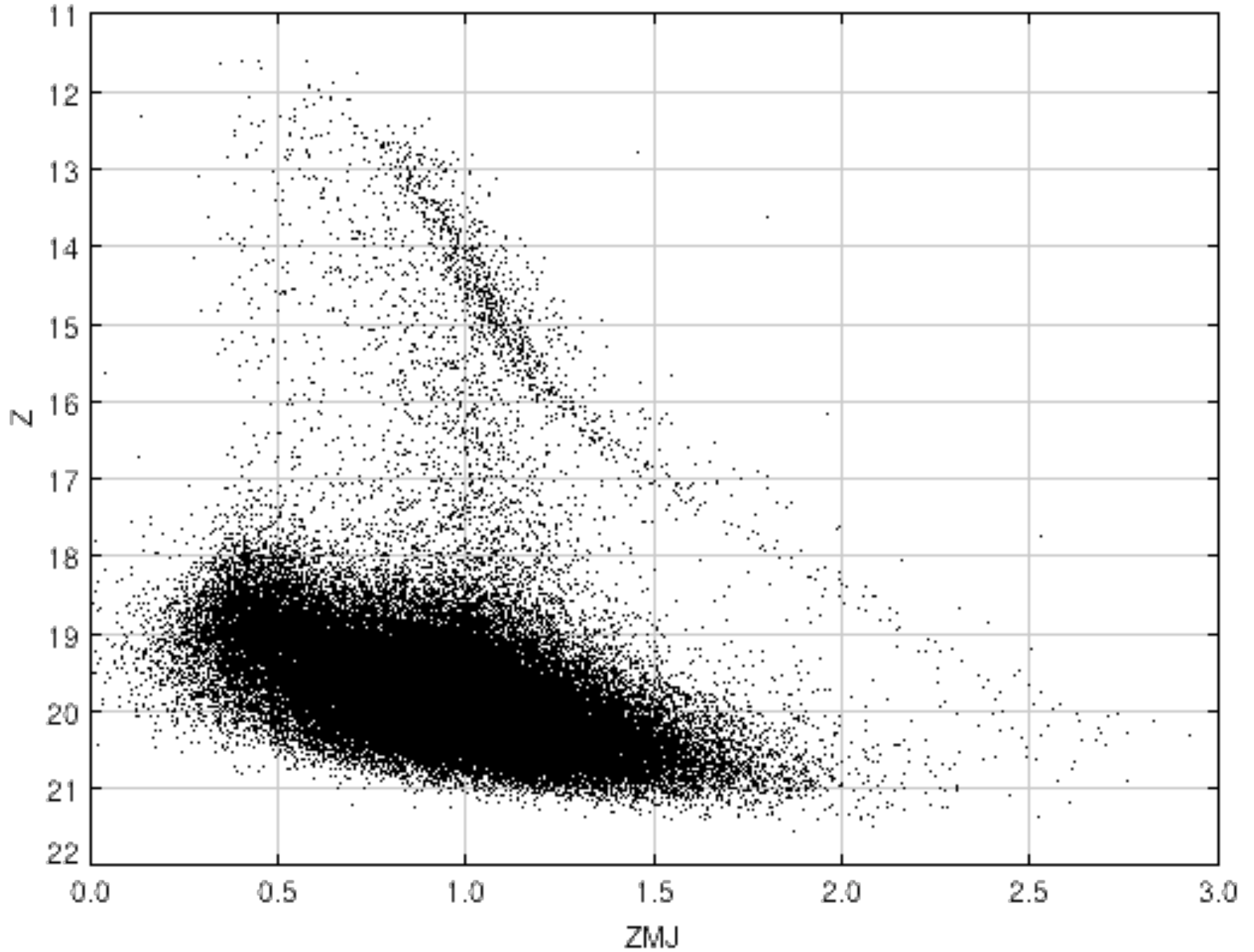
New in DR9: proper motion selection



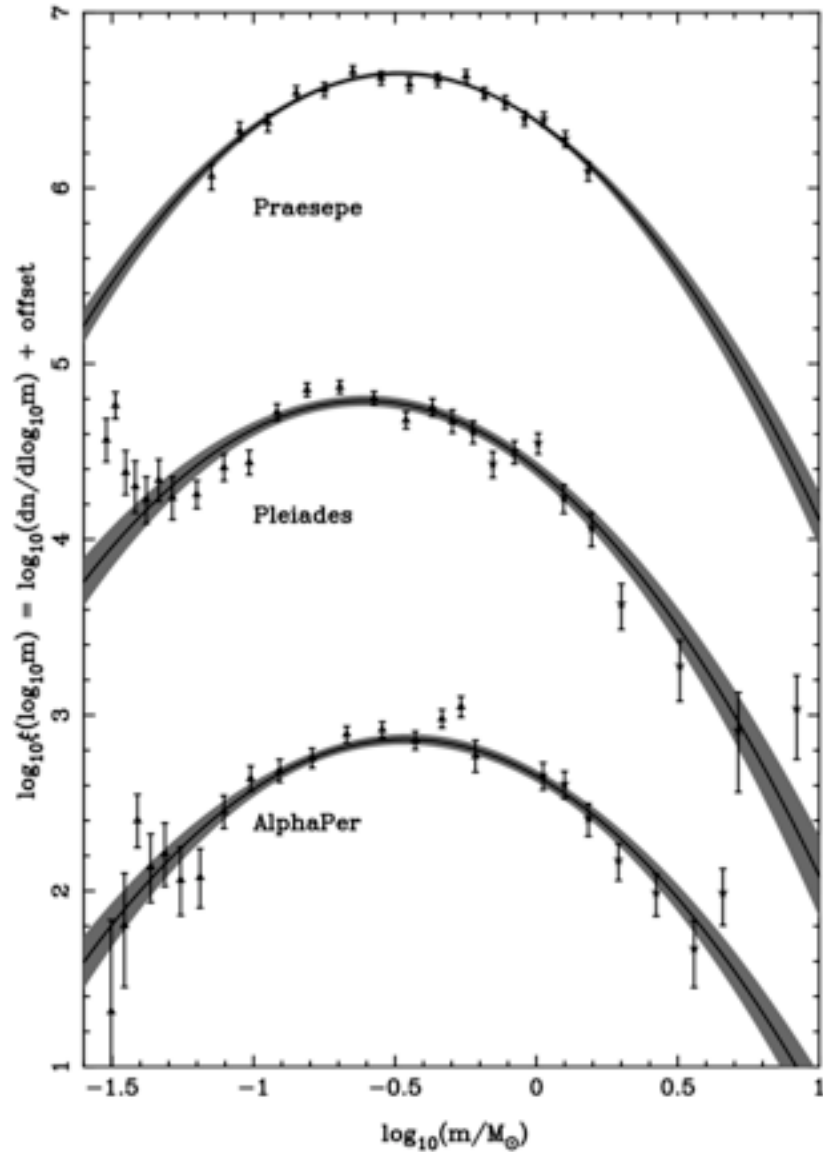
e.g. Praesepe

- Cluster members exhibit distinct proper motion
 - Dispersion in proper motions as low as 4 mas/yr
-
- add in a predicate of the form
... AND ABS(muRa – muRa0) < 3*sigMuRa AND ABS(muDec – muDec0) < 3*sigMuDec

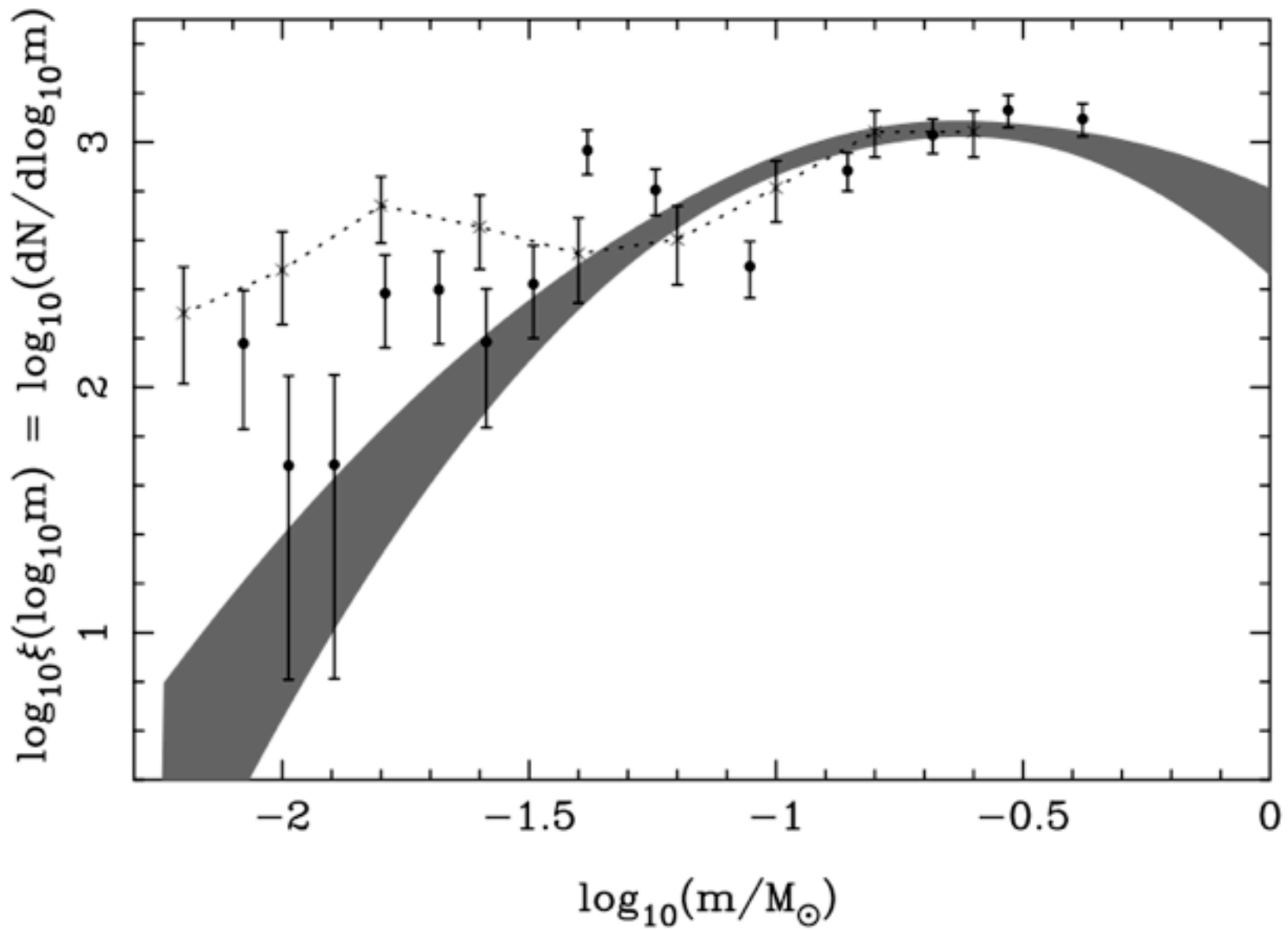


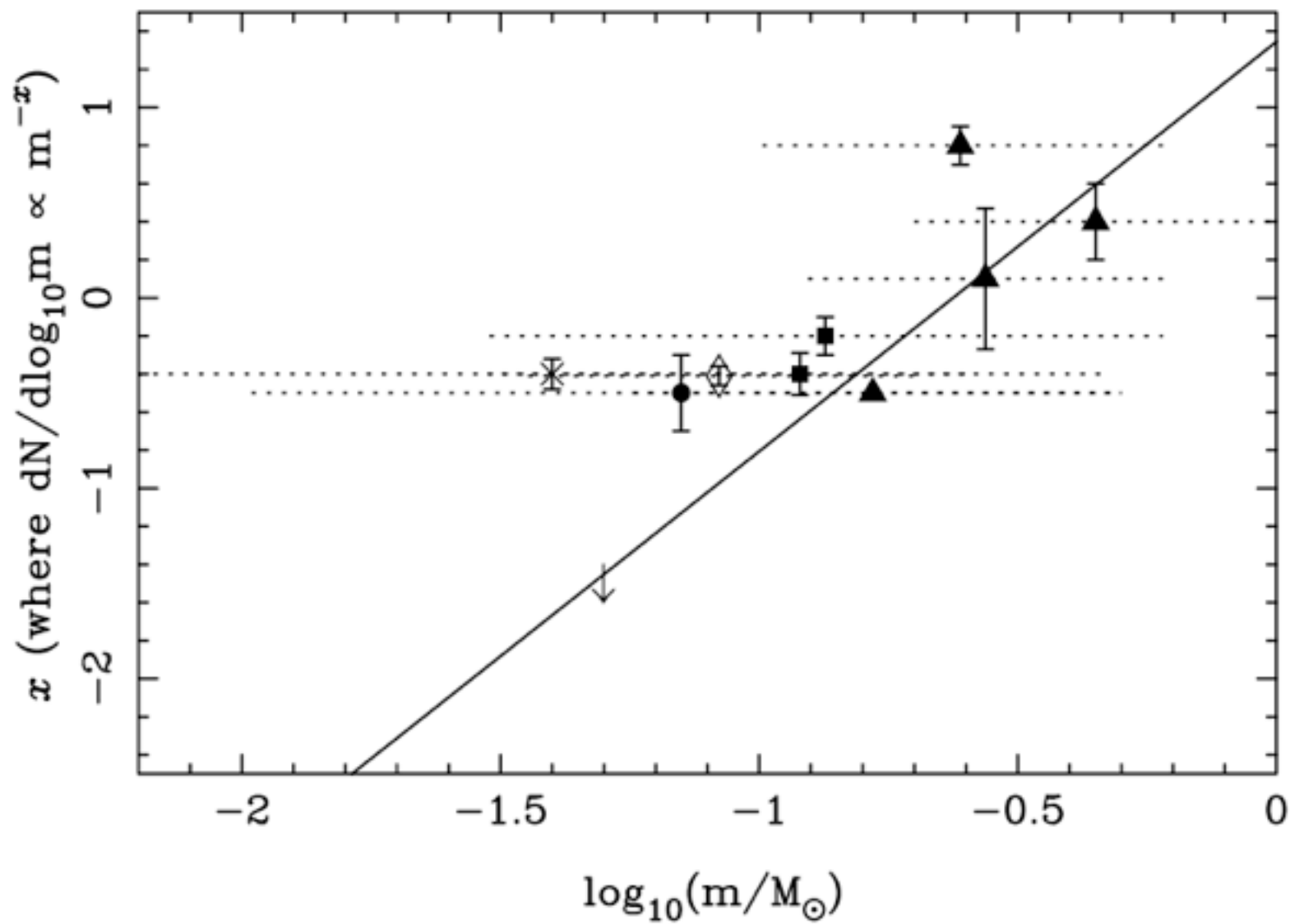


Evidence for MF variations...?



Population	Characteristic mass m_C (M_{\odot})	Dispersion σ	χ^2_{ν}
Alpha Per	0.344 ± 0.045	0.458 ± 0.019	2.275
Pleiades	0.247 ± 0.047	0.456 ± 0.023	4.382
Praesepe	0.328 ± 0.035	0.434 ± 0.015	0.962
Field (Chabrier 2003)	0.22	0.57	
Field (Chabrier 2005)	0.25	0.55	





Conclusions

- 5 colour, multi-epoch GCS imaging provides unrivalled datasets for VLMS/BD studies
- Survey completeness is good
- Preliminary evidence for MF variations, but need more sophisticated analysis

To find out more, type “Lodieu, N.” into ADS.