



Quantum Mechanics 3 2001/2002

Solution set 1

(1)

- (a) Must have $\psi(L) = \psi(-L) = 0$, so $B \cos kL + A \sin kL = 0$ and $B \cos kL - A \sin kL = 0$. Add or subtract these, to get $B = 0$ and $\sin kL = 0$, or $A = 0$ and $\cos kL = 0$.
- (b) This is easier, since $\psi(0) = 0$, so the \cos solution is never involved and $\psi = A \sin kx$, with $k = n\pi/(2L)$, $n = 1, 2, 3, \dots$
- (c) We need $\int_{-L}^L |A|^2 \cos^2 kL dx = 1$ for odd n . Since $\cos^2 y = (1 + \cos 2y)/2$, and the \cos part averages to zero, this gives $|A| = 1/L^{1/2}$ (the phase of A is arbitrary). The working for even n is similar.
- (d) By symmetry, $\langle x \rangle = 0$ (if we use the symmetric convention, where the well runs from $-L$ to L). The rms (or standard deviation) is the square root of the variance, which is $\langle x^2 \rangle - \langle x \rangle^2$; for the symmetric convention, we don't need the second term. For the second moment, we need $\langle x^2 \rangle = (1/L) \int_{-L}^L x^2 \sin^2(n\pi x/2L) dx$ (even n), or the same with $\sin \rightarrow \cos$ (odd n). Integration by parts gives

$$\langle x^2 \rangle = \frac{L^2}{3} \left(1 - \frac{6}{n^2 \pi^2} \right)$$

for any n . For the classical limit, $\langle x^2 \rangle = (1/2L) \int_{-L}^L x^2 dx$.

- (e) Inside the well, $V = 0$, so $E = \hbar^2 k^2 / 2m$, with k given by the boundary conditions.
- (f) For the second part, set $\hbar\omega$ ($= 2\pi\hbar c/\lambda$) to the difference in energy between the $n = 1$ and $n = 2$ states ($= 3\pi^2\hbar^2/[8mL^2]$). The effective size is therefore

$$L = \sqrt{\frac{3\pi\hbar\lambda}{16mc}} \simeq 1.66 \times 10^{-10} \text{ m.}$$

(2)

(a) Stationary state: $\psi(x, t) = u(x)e^{-i\omega t}$, where $\omega = E/\hbar$. Similarly, $\psi^*(x, t) = u^*(x)e^{+i\omega t}$. Therefore, the $e^{\pm i\omega t}$ terms cancel in $|\psi|^2 = \psi\psi^*$, leaving ρ independent of t . The same happens with \mathbf{j} .

If $\dot{\rho} = 0$, the given equation of continuity says $\nabla \cdot \mathbf{j} = 0$. This doesn't require \mathbf{j} to be a constant vector, but in 1D this is just $dj/dx = 0$, so j is constant.

(b) For a plane wave, $\nabla\psi = i\mathbf{k}\psi$. Also, $\nabla\psi^* = (\nabla\psi)^* = -i\mathbf{k}\psi^*$. Thus $\mathbf{j} = (\hbar/m)\mathbf{k}|\psi|^2$. Now, $\hbar k/m$ is momentum over mass, or velocity. So, j is probability density times velocity, reasonably enough if j is to be a flux density of probability.